

# Integral-equation approach to resonances in circular two-layer flows around an island with bottom topography

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This paper presents an integral-equation approach to the linear instability problem of two-layer quasi-geostrophic flows around circular islands with bottom topography. The study extends an earlier barotropic model of similar geometry and topography and focuses on the degree to which the topographic waves in the lower layer resonate with the basic flow in each layer. The integral approach poses the instability problem in a physically elucidating way, whereby the resonating neutral waves in the system are directly identified. The flows investigated are composed of uniform potential-vorticity (PV) rings in each layer, with the PV of each ring being of opposite sign. Four types of instabilities are identified: instability caused by the resonance of Rossby waves traveling along the liquid contours at the edge of each PV ring (CC resonance), instability caused by the resonance of the wave at the upper-layer contour and the topographic waves outside the lower-layer contour ( $C_1T$ ), a similar resonance of the lower-layer contour with the topographic waves ( $C_2T$ ), and a resonance between one eigenmode of the contour subsystem with the topographic waves (CCT). The three latter resonances lead to critical layer instabilities and can be identified as resonances between the contour waves and a collection of singular topographic modes with a critical layer. The  $C_1T$  ( $C_2T$ ) instability occurs when the lower-layer ring is sufficiently thin and the basic flow travels counterclockwise (clockwise). The neutral PV perturbations in the outer region behave asymptotically as barotropic (BT) or baroclinic (BC) modes that, when traveling clockwise, have spiral shapes and are wavelike in the radial direction. Usually, the BT mode is the mode in resonance with the contours but, for small growth rates, the BC mode may be the dominant mode. The nonlinear evolution of the CC resonance usually leads to emission of dipolar modons that then return to the island and are re-emitted in a quasiperiodic manner. The contour-topography instabilities may produce a narrow PV ring at the lower layer at the location of the critical layers of the dominant resonating topographic perturbations; this ring interacts with the original rings to form a quasi-stationary structure (e.g., a tripole) that rotates counterclockwise for a relatively long time before splitting into emitted modons.

## 1. Introduction

Islands in the stratified ocean can be surrounded by complex and variable current circulation patterns (?). Closed flows that are anomalous (i.e., flow in a direction opposite the overall circulation of the surrounding ocean) have been observed around Iceland, Taiwan, the islands of Kuril Chain (?) and the Pribilof Islands (?). In most cases, these anomalous circulations are anticyclonic (clockwise) in the northern hemisphere and are wind-driven. Waves generated in the vicinity of the islands may be trapped by the sloping topography or the coast and also contribute to the circulation strength; trapping by islands was shown for barotropic (BT) flows (see, e.g., Refs. ???) and for the stratified

sea (see, e.g., Refs. ???). In this paper, we study the conditions for instability of such baroclinic (BC) flows in an idealized model, where the island is circular and the bottom topography is conical (i.e., the beta-cone model).

To account for the stratification of the ocean, we use the simplified two-layer quasi-geostrophic (QG) model (?). The linear BC instability problem was solved by ? for the case of two-layer zonal uniform QG flow over a flat bottom and later solved by ? with the bottom topography included. For circularly symmetric flows, the instability problem was investigated by ? for BC QG vortices with a flat bottom and continuous stratification. Circular BC two-layer flows were investigated by ? for flows confined to an annular channel. Our model differs in several respects: First, it requires no external boundary. Second, the basic flow differs, as described below; in particular, currents in the two layers may flow in opposite directions in our model. Related to this is the fact that, in our case, the ratio of the bottom slope to the basic isopycnal slope is not constant, so this ratio plays no fundamental role.

? investigated the instability associated with idealized circularly symmetric BT currents around circular islands with bottom topography. There the flow was composed of two constant-potential-vorticity (constant-PV) rings around the island and the velocity outside the rings was zero. The purpose of this paper is to consider a variant of that model appropriate for a two-layer flow; now any layer consists of a single constant-PV ring. The flows in the two layers may have opposite directions and the velocity outside the rings does not vanish identically, but rather only the BT velocity. Figure 1 presents schematically the velocities and PV profiles.

The concept of resonance provides a physical interpretation of instabilities in two-layer shallow-water flows. As shown by many authors, different types of instabilities can be identified as resonances between neutral waves; the type of instability is determined by the interacting waves. The resonance usually occurs at the intersection of the dispersion curves [i.e., the curve of the phase velocity versus the wave number of the neutral modes (?)]. This was demonstrated in zonal shear flows (???) and zonal two-layer flows (????). In all these papers the resonance viewpoint was only applied to the case of shallow-water systems without the QG approximation; in this paper it is also applied to QG flows.

We identify the resonances in a simple way by using an integral-equation approach to the linear instability problem. To date, the integral approach has only been used for BT flows by ?, with no further extension elsewhere to BC flow. This paper fills this gap by showing that the integral approach poses the instability problem in a physically elucidating way whereby the coupling between the various wave types is directly identified.

The basic flow considered herein differs fundamentally from the BT case studied by ? because, here, the basic velocity is zero outside the rings in the lower layer, as is the PV gradient (see Figure 1). This fact permits the existence of singular neutral perturbations whose phase velocity equals the basic velocity at some distance [i.e., the lower layer has a critical layer (??)]. If a linear stability analysis shows that critical-layer eigenfunctions are neutrally stable, a more careful analysis of the initial-value problem would show that the eigenfunctions actually lead asymptotically over time to an algebraic time dependence (?). The time dependence is found mathematically from the singularities of the modes on the complex-frequency plane. Here on the beta cone is shown that new singularities, not present in the zonal case, appear; their damping effect is calculated analytically.

A critical layer instability (?) was observed experimentally by ? for a columnar vortex in stratified fluid and was studied in shallow-water single-layer flows (?). ? showed that this instability could be interpreted as the resonance between a nonsingular mode and a collection of singular modes. For the basic flow considered in this paper, resonances involving the topographic waves at the lower layer lead to a critical-layer instability.

Figure 1: Schematic profiles of the basic velocity in the upper layer  $\bar{V}_1$  (solid, red online) and in the lower layer  $\bar{V}_2$  (dashed, red online), and the PV in the upper layer  $Q_1$  (solid, blue online) and in the lower layer  $Q_2$  (dashed, blue online).

The resonating perturbations in this case are identified, and their effect on the nonlinear evolution of the flow is studied numerically.

The outline of this paper is as follows: §2 present the basic equations of the model of quasigeostrophic two-layer flows, and §3 gives the derivation of the integral eigenvalue equation of the linear stability analysis. §4 applies the integral equation to the basic flow composed of the two-layer PV rings plotted schematically in Figure 1. The resonance viewpoint is then presented for this flow in §5. §6 discusses the spectrum of solutions in the exterior region  $r > R_2$ . These neutrally stable solutions (according to a linear stability analysis) may resonate with the waves at the contours of the PV discontinuities at  $R_1$  and  $R_2$ , and their time-dependent damping is also found. §7 further explores the resonances via dispersion curves, calculates the growth rates and structure of unstable perturbations, and discusses the conditions for the dominance of BT versus BC couplings. Finally, §8 examines how instability type affects the nonlinear evolution of the flow.

## 2. Two-layer flows on a beta cone: Governing equations

Consider a two-layer QG model in which the flow surrounds a cylindrical island. The bottom around the island is assumed to have a constant radial slope so that the depth increases linearly with distance from the island. Under the QG approximation and the rigid-lid condition at the sea surface, the flow is effectively two dimensional in each layer. The variables of the upper and lower layers are denoted by the subscripts 1 and 2, respectively. The unperturbed layer thickness is denoted  $H_i$  ( $i = 1, 2$ ), and the sum of layers by  $H$ . In the polar coordinates  $r$  and  $\theta$ , the radial and azimuthal components of the velocity,  $u_i$  and  $v_i$ , respectively, in each layer ( $i = 1, 2$ ) can be expressed in terms of a streamfunction  $\Psi_i$  by

$$u_i = -\frac{1}{r} \frac{\partial \Psi_i}{\partial \theta}, \quad v_i = \frac{\partial \Psi_i}{\partial r}. \quad (2.1)$$

In the following, whenever the subscript  $i$  appears, it refers to layer  $i$ . The slope at the bottom introduces a linear term in  $r$  for the PV at the lower layer (see ? for details). The proportionality constant is the topographic beta,  $\beta = -f \tan(\alpha)/H_2$ , where  $f$  is the Coriolis parameter. The island is assumed to be small compared with the planetary scale, so  $f$  may be regarded as constant (this is analogous to the  $f$ -plane approximation, cf. ?). For an island in the northern hemisphere,  $\beta$  is negative.

In terms of the streamfunctions, the PVs in layers 1 and 2 are defined as (cf. ?)

$$Q_1 = \nabla^2 \Psi_1 - \frac{f^2}{g'H_1} (\Psi_1 - \Psi_2), \quad Q_2 = \nabla^2 \Psi_2 + \frac{f^2}{g'H_2} (\Psi_1 - \Psi_2) + \beta r, \quad (2.2)$$

where  $g' = g(\rho_2 - \rho_1)/\rho_1$  is the reduced gravity ( $g$  being the acceleration due to gravity, and  $\rho_1$  and  $\rho_2$  being the layer densities).

On the beta cone, a natural length scale is the radius  $R$  of the island. We are interested in flows whose horizontal length scale is  $R$ , such that the curvature plays a dominant role, so  $r \approx R$ . Flows having much smaller length scale behave locally as straight flows, whereas at much larger length scales the island's influence is negligible. In §3 the basic flow is defined, where the PV in the upper layer is constant inside a ring,  $\Gamma_1$ . This determines the timescale  $1/|\Gamma_1|$ . Therefore, assuming that time scales advectively, we

transform the variables into nondimensional variables via

$$t \rightarrow t/|\Gamma_1|, \quad r \rightarrow Rr, \quad Q_i \rightarrow |\Gamma_1|Q_i, \quad \Psi_i \rightarrow |\Gamma_1|R^2\Psi_i, \quad \beta \rightarrow |\Gamma_1|\beta/R. \quad (2.3)$$

Nondimensionalization of equations (2.2) then yields

$$Q_1 = \nabla^2\Psi_1 - \frac{\Lambda^2}{\lambda_1}(\Psi_1 - \Psi_2), \quad Q_2 = \nabla^2\Psi_2 + \frac{\Lambda^2}{\lambda_2}(\Psi_1 - \Psi_2) + \beta r, \quad (2.4)$$

where  $\Lambda^2 = (R/L_{Ro})^2$  is the reverse Burger number, and  $L_{Ro} = (g'H/f_0^2)^{1/2}$  is the Rossby deformation radius. In the ocean,  $L_{Ro}$  varies from about 1 km at high latitudes to about 400 km at the equator (?). Small islands may have a radius of few kilometers, whereas large islands may have a radius of 200 km. Therefore,  $\Lambda$  may change from  $10^{-4}$  to 200. To be consistent with the QG approximation presented above,  $\Lambda$  should be of order unity or less (?). Therefore, we mostly use  $\Lambda = 1$ ; this means that the island's size is of the same order of magnitude as the Rossby deformation radius. The relative thickness of each layer  $i$  is denoted  $\lambda_i = H_i/H$  ( $i = 1, 2$ ), with  $H$  being the total thickness of the fluid:  $H = H_1 + H_2$ .

The PV conservation equations governing the dynamics are

$$\frac{\partial Q_i}{\partial t} + \frac{1}{r} \left( \frac{\partial \Psi_i}{\partial r} \frac{\partial Q_i}{\partial \theta} - \frac{\partial \Psi_i}{\partial \theta} \frac{\partial Q_i}{\partial r} \right) = 0 \quad (i = 1, 2). \quad (2.5)$$

### 3. The integral eigenvalue equations

We represent the PVs  $Q_1$  and  $Q_2$  and the streamfunctions  $\Psi_1$  and  $\Psi_2$  of the flow as sums of the basic-state values (indicated by a bar) and the associated perturbations,

$$Q_i = \bar{Q}_i(r) + q_i(r, \theta, t), \quad \Psi_i = \bar{\Psi}_i(r) + \psi_i(r, \theta, t). \quad (3.1)$$

Assuming small perturbations, the linearized equations expressing conservation of PV that result from (2.1) and (2.5) are

$$\frac{\partial q_1}{\partial t} + \frac{\bar{V}_1}{r} \frac{\partial q_1}{\partial \theta} - \frac{1}{r} \frac{\partial \psi_1}{\partial \theta} \frac{d\bar{Q}_1}{dr} = 0, \quad \frac{\partial q_2}{\partial t} + \frac{\bar{V}_2}{r} \frac{\partial q_2}{\partial \theta} - \frac{1}{r} \frac{\partial \psi_2}{\partial \theta} \frac{d\bar{Q}_2}{dr} = 0. \quad (3.2)$$

The perturbations are considered to be associated with an azimuthal integer mode number  $m$  and (generally complex) frequency  $\omega$ :

$$\{q_i(r, \theta, t), \psi_i(r, \theta, t)\} = \{Q_i(r), \Phi_i(r)\}e^{i(m\theta - \omega t)}, \quad (3.3)$$

where we suppress the explicit notation of  $m$  in  $Q_i(r)$  and  $\Phi_i(r)$  to simplify the notation; the notation  $m$  is also dropped in subsequent expressions. Using (3.3) in (3.2) yields the Rayleigh equations,

$$\left( \frac{\bar{V}_i(r)}{r} - \frac{\omega}{m} \right) Q_i - \frac{\Phi_i}{r} \frac{d\bar{Q}_i}{dr} = 0. \quad (3.4)$$

By using (2.4) and (3.3), the functions  $Q_i(r)$  and  $\Phi_i(r)$  are related via the equations

$$Q_1 = \frac{d^2\Phi_1}{dr^2} + \frac{1}{r} \frac{d\Phi_1}{dr} - \frac{m^2}{r^2}\Phi_1 - \frac{\Lambda^2}{\lambda_1}(\Phi_1 - \Phi_2), \quad (3.5)$$

$$Q_2 = \frac{d^2\Phi_2}{dr^2} + \frac{1}{r} \frac{d\Phi_2}{dr} - \frac{m^2}{r^2}\Phi_2 + \frac{\Lambda^2}{\lambda_2}(\Phi_1 - \Phi_2). \quad (3.6)$$

Given  $Q_1$  and  $Q_2$ , equations (3.5) and (2) for the streamfunctions can be decoupled. Because the term  $\beta r$  is absent, the decoupling is possible here, in contrast with equations

(2.4) where it is not possible because equations (3.5) and (2) deal with the perturbations of the PVs. Consider the BT and BC stream-function perturbations

$$\Phi_{\text{BT}} = \lambda_1 \Phi_1 + \lambda_2 \Phi_2, \quad \Phi_{\text{BC}} = \Phi_1 - \Phi_2, \quad (3.7)$$

and the corresponding PV perturbations

$$\mathcal{Q}_{\text{BT}} = \lambda_1 \mathcal{Q}_1 + \lambda_2 \mathcal{Q}_2, \quad \mathcal{Q}_{\text{BC}} = \mathcal{Q}_1 - \mathcal{Q}_2. \quad (3.8)$$

From equations (3.5) and (2) and the definitions (3.7) and (3.8), we get the equations

$$\frac{d^2 \Phi_{\text{BT}}}{dr^2} + \frac{1}{r} \frac{d\Phi_{\text{BT}}}{dr} - \frac{m^2}{r^2} \Phi_{\text{BT}} = \mathcal{Q}_{\text{BT}}, \quad (3.9)$$

$$\frac{d^2 \Phi_{\text{BC}}}{dr^2} + \frac{1}{r} \frac{d\Phi_{\text{BC}}}{dr} - \frac{m^2}{r^2} \Phi_{\text{BC}} - \frac{\Lambda^2}{\lambda_1 \lambda_2} \Phi_{\text{BC}} = \mathcal{Q}_{\text{BC}}, \quad (3.10)$$

where  $\lambda_1 + \lambda_2 = 1$  was used in the last equation.

The general solutions to (3.9) and (3.10) can be written as

$$\Phi_{\text{BT}}(r) = \int_R^\infty G_{\text{BT}}(r, r') \mathcal{Q}_{\text{BT}}(r') dr', \quad \Phi_{\text{BC}}(r) = \int_R^\infty G_{\text{BC}}(r, r') \mathcal{Q}_{\text{BC}}(r') dr', \quad (3.11)$$

where  $G_{\text{BT}}(r, r')$  and  $G_{\text{BC}}(r, r')$  are the BT and BC Green's functions, respectively. The derivations and expressions for these Green's functions appear in Appendix B. From (3.7) we get the expression of the streamfunction in each layer in terms of the BT and BC modes,

$$\Phi_1 = \Phi_{\text{BT}} + \lambda_2 \Phi_{\text{BC}}, \quad \Phi_2 = \Phi_{\text{BT}} - \lambda_1 \Phi_{\text{BC}}. \quad (3.12)$$

Using (3.8), (3.11), and (3.12) we get

$$\Phi_1(r) = \int_R^\infty [G_{11}(r, r') \mathcal{Q}_1(r') + G_{12}(r, r') \mathcal{Q}_2(r')] dr', \quad (3.13)$$

$$\Phi_2(r) = \int_R^\infty [G_{21}(r, r') \mathcal{Q}_1(r') + G_{22}(r, r') \mathcal{Q}_2(r')] dr', \quad (3.14)$$

where the four Green's functions  $G_{ij}$  ( $i, j = 1, 2$ ) are defined as

$$G_{11}(r, r') = \lambda_1 G_{\text{BT}}(r, r') + \lambda_2 G_{\text{BC}}(r, r'), \quad (3.15)$$

$$G_{12}(r, r') = \lambda_2 [G_{\text{BT}}(r, r') - G_{\text{BC}}(r, r')], \quad (3.16)$$

$$G_{21}(r, r') = \lambda_1 [G_{\text{BT}}(r, r') - G_{\text{BC}}(r, r')], \quad (3.17)$$

$$G_{22}(r, r') = \lambda_2 G_{\text{BT}}(r, r') + \lambda_1 G_{\text{BC}}(r, r'). \quad (3.18)$$

Based on (3.13) and (3.14), the function  $G_{ij}$  is the Green's function that connects a PV perturbation in layer  $j$  to the streamfunction in layer  $i$ . Note also that the no-slip boundary condition at the cylindrical wall [i.e.,  $\Phi_1(R) = \Phi_2(R) = 0$  by (2.1) and (3.3)] is satisfied automatically by (3.13) and (3.14). We now express the streamfunctions in terms of the PV perturbations by inserting (3.13) and (3.14) into (3.4) to get

$$\frac{m \bar{V}_i(r)}{r} \mathcal{Q}_i - \frac{m}{r} \frac{d\bar{Q}_i(r)}{dr} \int_R^\infty [G_{i1}(r, r') \mathcal{Q}_1 + G_{i2}(r, r') \mathcal{Q}_2] dr' = \omega \mathcal{Q}_i(r). \quad (3.19)$$

Equation (3.19) constitutes a system of two linear integral equations for the PV perturbations at both layers. In the next section, we apply these equations to flows composed of two-layer constant-PV rings.

## 4. Flows composed of two-layer rings

### 4.1. Basic flow profile

As stated above, in the subsequent stability analysis, we take as a basic state a circularly symmetric flow composed of a uniform-PV ring in each layer. The ring in the upper (lower) layer is bounded by a rigid contour at  $r = R$  and a material contour at  $r = R_1$  ( $R_2$ ), the latter of which we denote  $C_1$  ( $C_2$ ). Outside the rings, the PV of each layer equals the background PV. Denoting by  $\bar{Q}_i$  the PV of the basic flow and by  $\Gamma_1$  and  $\Gamma_2$  the PV in the upper and lower rings, respectively, we write

$$\bar{Q}_1(r) = \begin{cases} \Gamma_1, & R \leq r \leq R_1 \\ 0, & R_1 < r, \end{cases} \quad \bar{Q}_2(r) = \begin{cases} \Gamma_2, & R \leq r \leq R_2 \\ \beta r, & R_2 < r. \end{cases} \quad (4.1)$$

The PV-jumps across each contour are

$$\Delta_1 = -\Gamma_1, \quad \Delta_2 = \beta R_2 - \Gamma_2. \quad (4.2)$$

The expressions for the basic streamfunctions  $\bar{\Psi}_i$  and velocities  $\bar{V}_i$  resulting from this PV configuration are derived in Appendix A. Since the flow is attached to a rigid cylindrical wall (the island), a natural (although unnecessary) boundary condition would be the no-slip condition (i.e., the vanishing of the velocity at  $r = R$ ). This condition results from the role of turbulent viscosity in the vicinity of the vertical wall during the formation of the closed flow, as explained in detail in Ref. ?. The vanishing of the velocity at the rigid boundary at  $r = R$  imposes the following relation between  $\Gamma_1$  and  $\Gamma_2$  (see Appendix A for details):

$$\Gamma_2 = \frac{-2R^3\beta\lambda_2 + 2R_2^3\beta\lambda_2 + 3\Gamma_1R^2\lambda_1 - 3\Gamma_1R_1^2\lambda_1}{3\lambda_2(R_2^2 - R^2)}. \quad (4.3)$$

Figure 1 shows schematic profiles of the velocities and PVs in both layers.

For future reference, we note that (4.1) may be written as

$$\bar{Q}_1(r) = \Gamma_1 + \Delta_1\mathcal{H}(r - R_1), \quad \bar{Q}_2(r) = \Gamma_2 + (\beta r - \Gamma_2)\mathcal{H}(r - R_2), \quad (4.4)$$

where  $\mathcal{H}(\cdot)$  is the Heaviside function, which is defined to vanish at zero:  $\mathcal{H}(0) = 0$ . The gradient of the basic PV profile (4.4) is

$$\frac{d\bar{Q}_1}{dr} = \Delta_1\delta(r - R_1), \quad \frac{d\bar{Q}_2}{dr} = \Delta_2\delta(r - R_2) + \beta\mathcal{H}(r - R_2). \quad (4.5)$$

### 4.2. Integral eigenvalue equations

Define  $s_i(r, \theta, t)$  to be the displacement of a particle from its initial reference location at  $t = 0$  in layer  $i$ . Since PV is conserved as it moves, the change in PV at the new particle location is, for small  $s_i$ ,

$$q_i(r + s_i, \theta, t) = \bar{Q}_i(r, \theta) - \bar{Q}_i(r + s_i, \theta) = -\frac{d\bar{Q}_i}{dr}s_i \quad (4.6)$$

(cf. Ref. ?). If all displacements are associated with an azimuthal integer mode number  $m$  and frequency  $\omega$  as in (3.3), then we may write  $s_i = d_i(r)e^{i(m\theta - \omega t)}$  where  $d_i(r)$  is the amplitude of the radial displacement of the particle. Comparing (4.6) with (3.3) shows clearly that

$$\mathcal{Q}_i(r) = -\frac{d\bar{Q}_i}{dr}d_i(r). \quad (4.7)$$

Based on (4.5) and (4.7),  $\mathcal{Q}_1$  vanishes everywhere except at  $r = R_1$ , where it is given by a delta function. The displacement  $s_1$  of a particle at  $r = R_1$  can also be interpreted

as the deformation of  $C_1$  (cf. Refs. ??); the amplitude  $d_1(R_1)$  of the perturbation  $C_1$  is denoted by  $\alpha_1/R_1$ . Similarly, the amplitude  $d_2(R_2)$  of perturbation  $C_2$  at  $r = R_2$  is denoted by  $\alpha_2/R_2$ . The amplitude of the displacement  $d_2(r)$  at the outer region  $r > R_2$  in the lower layer is denoted by  $\eta(r)/r$  and can be interpreted as the deformation of the background constant-PV contours (which are circles). The division by  $R_1$ ,  $R_2$ , and  $r$  is done to make the integral operator symmetric, which is useful as is shown below (§6). Therefore, using (4.5) and (4.7), we write

$$\mathcal{Q}_1 = -\frac{\Delta_1 \alpha_1}{R_1} \delta(r - R_1), \quad \mathcal{Q}_2 = -\frac{\Delta_2 \alpha_2}{R_2} \delta(r - R_2) - \frac{\beta}{r} \eta(r) \mathcal{H}(r - R_2). \quad (4.8)$$

Inserting (4.8) into the eigenvalue integral equations (3.19) yields the following three eigenvalue equations:

$$\frac{\bar{V}_1(R_1) - \Delta_1 G_{11}(R_1, R_1)}{R_1} \alpha_1 - \frac{\Delta_1 G_{12}(R_1, R_2)}{R_2} \alpha_2 - \beta \Delta_1 \int_{R_2}^{\infty} \frac{G_{12}(R_1, r')}{r'} \eta(r') dr' = \frac{\omega}{m} \alpha_1, \quad (4.9)$$

$$-\frac{\Delta_2 G_{21}(R_2, R_1)}{R_1} \alpha_1 + \frac{\bar{V}_2(R_2) - \Delta_2 G_{22}(R_2, R_2)}{R_2} \alpha_2 - \beta \Delta_2 \int_{R_2}^{\infty} \frac{G_{22}(R_2, r')}{r'} \eta(r') dr' = \frac{\omega}{m} \alpha_2, \quad (4.10)$$

$$-\frac{G_{21}(r, R_1)}{R_1} \alpha_1 - \frac{G_{22}(r, R_2)}{R_2} \alpha_2 + \frac{\bar{V}_2(r)}{r} \eta(r) - \beta \int_{R_2}^{\infty} \frac{G_{22}(r, r')}{r'} \eta(r') dr' = \frac{\omega}{m} \eta(r). \quad (4.11)$$

Equations (4.9)–(4.11) can be recast into a standard matrix eigenvalue equation, which can then be solved numerically by using the “eig” function in Matlab, which uses the QZ algorithm (?). To get the matrix form, the integrals are approximated via a Gaussian quadrature rule (see, e.g., Ref. ?), which for any function  $f(x)$  takes the form  $\int_{R_2}^{\infty} f(r) dr \approx \sum_{i=1}^N w_i f(r_i)$ ; here  $r_i$  and  $w_i$  are the nodes and weights, respectively, of the quadrature rule employed.

Since the domain is infinite, we divide the integral into two regions: The first region is the near neighborhood of the island, where the basic flow velocities at the two layers are significant. Outside the largest ring, the velocities drop exponentially with  $r$  with typical length scale  $\Lambda/\sqrt{\lambda_1 \lambda_2}$  (see Appendix A); thus the velocities remain significant at  $R_2 \leq r \leq \max(R_1, R_2) + 5\Lambda/\sqrt{\lambda_1 \lambda_2}$ . In this region, the Legendre-Gauss quadrature rule is applied with 1000 points. The second region is outside, at  $\max(R_1, R_2) + 5\Lambda/\sqrt{\lambda_1 \lambda_2} < r < \infty$ , where the Gauss-Laguerre quadrature rule is applied with 150 points. Tests of convergence show that the results are robust; e.g., even with half the number of points in each region, the error in calculating the eigenvalues is less than 0.1%.

## 5. Resonance viewpoint

The integral eigenvalue equations (4.9)–(4.11) allow direct interpretation of the couplings that occur in the system studied. We demonstrate it by using the first equation, (4.9), which determines the angular velocity of the upper contour  $C_1$  perturbation, at  $r = R_1$ . The right-hand side (RHS) of (4.9) may also be viewed as the time derivative of the PV perturbation at  $r = R_1$  because the time derivative is proportional to  $\omega$  [cf. (3.3)]. The first term on the left-hand side (LHS) of (4.9) contains the free-streaming term  $\bar{V}_1(R_1)\alpha_1/R_1$  with coupling to the basic PV jump at its place, namely,  $-\Delta_1 G_{11}(R_1, R_1)\alpha_1/R_1$ . This term would determine the angular velocity of the PV con-

Figure 2: Growth rates  $\text{Im}(\omega)$  for different couplings as functions of  $R_2$  at  $R_1 = 5$ ,  $\beta = -0.1$ ,  $m = 2$ ,  $\Lambda = 1$ , and  $\lambda_1 = \lambda_2 = 0.5$  for (a)  $\Gamma_1 = 1$ , (b)  $\Gamma_1 = -1$ . The types of couplings are as follows: Full (red), CC (blue), C1T (purple), C2T (brown), CCT (green).

tour at  $r = R_1$  were no other couplings to occur (cf. Ref. ?). The next term represents the coupling between the PV perturbations at  $C_1$  and  $C_2$  because  $\alpha_2$  is what influences the time dependence of  $\alpha_1$ . Finally, the integral term dictates how the PV perturbation  $\eta(r)/r$  in the lower layer outside  $C_2$  affects the evolution of the perturbation at  $C_1$ .

Each of the coupling terms can now be identified when instability is reached. By allowing only certain couplings to remain in the equations while removing others, different subsystems of the entire system can be isolated and the dominant ones found. These are the couplings that lead to the closest phase velocity and growth rate of the fully coupled system. In this case, the PV perturbations that couple and thereby cause the instability are said to be resonant.

The resonance viewpoint has been used by many authors for shallow-water systems, as mentioned in the Introduction. Usually, resonance is demonstrated by the intersection of two dispersion curves (?); in the QG case, we find the dominant couplings in the eigenvalue equations written in terms of the PVs. §7 shows that the results are consistent with those obtained based on the intersection of two dispersion curves.

In the resonance viewpoint, the instability is caused by the interaction between two waves that phase-lock and thereby enhance each other's growth (see, e.g., Refs. ??). For the basic flow considered herein, three Rossby waves can interact: the first wave travels along  $C_1$  (where the PV in the upper layer jumps), the second wave travels along  $C_2$ , and the third wave, which exists because of the bottom topography, travels in the outer region ( $r > R_2$ ) of the second layer. As is discussed in §6, various types of perturbations exist for this third wave, which we collectively call “topographic” perturbations. By using the eigenvalue equations (4.9)–(4.11), we identify four types of instabilities:

- (a) Contour-contour (CC) instability: In this case, the dominant interaction leading to instability is that between perturbations at the ring periphery (i.e., between  $C_1$  and  $C_2$ ). The coupling contours  $C_1$  and  $C_2$  alone correspond to setting  $\beta = 0$  in (4.9)–(4.11), thus leaving two algebraic equations to be solved. §7 presents this instability in more detail.
- (b) Contour- $C_1$ -topography (C1T) instability: In this case the CC subsystem (composed of  $C_1$  and  $C_2$  alone) is stable; that is, the PV-jumps alone do not cause the instability, but rather the resonance of the wave at  $C_1$  with the topographic PV perturbations in the lower layer (in the region  $r > R_2$ ). The eigenvalue calculation in this case is done by setting  $\alpha_2 = 0$  in (4.9)–(4.11).
- (c) Contour- $C_2$ -topography (C2T) instability: In this case, the wave on the lower layer PV contour is in resonance with the topographic PV perturbations outside the contour. The eigenvalue calculation in this case is done by setting  $\alpha_1 = 0$  in (4.9)–(4.11).
- (d) Both-contours-topography (CCT) instability: In this case, the dominant resonance is between one of the neutral perturbation types of the mutual contours subsystem CC and the topographic perturbations. The eigenvalue calculation in this case is done by rearranging (4.9)–(4.11) such that the perturbations of the CC subsystem are decoupled; this is explained in §7.2. Note that, although it seems that the entire system takes part in this instability, this is not so: only one of the neutral CC perturbation types participates in this resonance, whereas the other one does not.



In the following, we collectively call instabilities (ii)–(iv) contour-topography (CT) instabilities; these are discussed in §7 in more detail. Figure 2 presents an example showing the identification of the types of instabilities; the growth rates [i.e.,  $\text{Im}(\omega)$ ] are shown for each of the above resonances as a function of the radius of the lower-layer ring. The flow parameters are  $\beta = -0.1$ ,  $R_1 = 5$ ,  $\lambda_1 = \lambda_2 = 1/2$ , and  $\Gamma_1 = 1$  (Figure 2a) or  $\Gamma_1 = -1$  (Figure 2b).

For  $\Gamma_1 = 1$  (Figure 2a), the CC resonance dominates when  $2.3 < R_2 < 5.5$ , whereas the  $C_2T$  resonance dominates when  $R_2 < 2$ . The  $C_1T$  resonance is completely absent in this case, as explained below (§5.1). The growth rates of the CC instability generally exceed those of the  $C_2T$  instability. Also, there is a small “window” at  $2 < r < 2.3$  where the CC interaction is stable whereas the  $C_2T$  interaction is not; however, the growth rate of the full instability is much greater (up to fourfold) than the growth rate of the  $C_2T$  resonance and therefore cannot be attributed to this resonance. The dominant instability in this region is of type CCT.

For  $\Gamma = -1$  (Figure 2b), again the CC resonance leads to the greatest growth rates and is dominant for most values of  $R_2$ . In much of the CC instability region, the actual (full-system) growth rate is less than that implied by the growth rate of the CC interaction. Therefore, the topography in this case stabilizes the flow. Again at small values of  $R_2$  (below 1.5), the dominant resonance is between one of the contours and the topographic perturbations, but this time it is of type  $C_1T$ , and the  $C_2T$  type is absent. At small regions  $1.5 < r < 2.1$ , the instability is again of type CCT.

### 5.1. Pseudomomentum considerations

Although momentum is not a conserved quantity in the system (3.2) of linearized equations, one can define an analogous quantity that is conserved, namely, the pseudomomentum (?). Whereas a necessary condition for instability to occur is phase-locking (i.e., the intersection of the dispersion curves of two neutral waves), not every intersection leads to instability. As shown in Ref. ?, an additional requirement for instability is that the two waves have pseudomomenta of opposite signs.

The expression for the pseudomomentum density in the two-layer model on the beta cone (i.e., in polar coordinates where the basic flow is radially symmetric) is developed in Appendix C and is given by

$$\mathcal{M} = -\frac{\lambda_1}{2} \frac{d\bar{Q}_1}{dr} \langle s_1^2 \rangle - \frac{\lambda_2}{2} \frac{d\bar{Q}_2}{dr} \langle s_2^2 \rangle, \quad (5.1)$$

where the brackets  $\langle \cdot \rangle$  denote the azimuthal average of the variable. The pseudomomentum density satisfies the continuity equation

$$\frac{\partial \mathcal{M}}{\partial t} + \frac{1}{r} \frac{\partial \mathcal{F}}{\partial r} = 0, \quad (5.2)$$

where  $\mathcal{F} = \langle \frac{\partial \psi_1}{\partial \theta} \frac{\partial \psi_1}{\partial r} \rangle$  is the Eliaseen-Palm flux. If (5.2) is integrated over the entire plane outside the island ( $r > R$ ), we get the equation for pseudomomentum conservation,  $\frac{\partial M}{\partial t} = 0$ , where

$$M = \int_R^\infty r \mathcal{M} dr = - \int_R^\infty \left( \frac{\lambda_1}{2} \frac{d\bar{Q}_1}{dr} \langle s_1^2 \rangle + \frac{\lambda_2}{2} \frac{d\bar{Q}_2}{dr} \langle s_2^2 \rangle \right) dr. \quad (5.3)$$

Since  $M$  is conserved, it must vanish in the case of an instability, which leads to Rayleigh’s necessary condition for instability that the basic PV gradient must be somewhere negative and somewhere positive (cf. ??). Moreover, if only two perturbation types are in resonance, their pseudomomenta must have opposite signs (?).

For the basic flow considered herein, the use of (4.5) and (4.8) gives a pseudomomentum of

$$M = \left( -\frac{\lambda_1 \Delta_1}{4R_1^2} |\alpha_1|^2 - \frac{\lambda_2 \Delta_2}{4R_2^2} |\alpha_2|^2 - \int_R^\infty \frac{\lambda_2 \beta}{4r^2} |\eta|^2 dr \right) e^{2\text{Im}(\omega)t}. \quad (5.4)$$

Since the PV jumps when the two contours  $C_1$  and  $C_2$  have opposite signs, their pseudomomenta have opposite signs and the Rossby waves traveling along these contours may be in resonance. The pseudomomentum of the perturbation at the exterior region  $r > R_2$  is always positive,  $\beta$  being always negative. Therefore, the exterior perturbations can only be in resonance with the contour wave whose pseudomomentum is negative (i.e., traveling along a positive PV gradient).

This explains why only one contour wave resonates with the topographic perturbations, as shown in Figure 2. If  $\Gamma_1 = 1$ , the pseudomomentum of the contour wave at  $r = R_1$  is positive [because  $\Delta_1 < 0$  by (4.2)], whereas the pseudomomentum of the contour wave at  $r = R_2$  is negative [because  $\Delta_2 > 0$  for the specific parameters dictated by (4.2)]. Thus, only the lower-layer contour has pseudomomentum opposite in sign relative to the outside perturbations (which always have positive pseudomomentum), so a  $C_1T$  instability is impossible in this case (Figure 2a). The same argument explains why a  $C_2T$  instability is impossible for  $\Gamma_1 = -1$  (Figure 2b).

## 6. Perturbation types in outer region

We now focus on the subsystem of the basic flow outside the liquid contours, which means that we search for modes whose perturbation is dominant (i.e., strong relative to the contours' perturbations) in the lower layer at  $r > R_2$ . For this, we assume that the PV-jumps at any of the liquid contours are negligible ( $\Delta_1 \approx \Delta_2 \approx 0$ ), thus avoiding any coupling to waves at the given contour. The resulting PV perturbations can be seen as self-excitations of the outer region caused by the topography. As a consequence, the contours may physically oscillate and resonate to yield the CT instability, which is discussed in §7.

Neglecting  $\alpha_1$  and  $\alpha_2$  in (4.9)–(4.11) yields a single integral equation for  $\eta$ ,

$$\frac{\bar{V}_2(r)}{r} \eta(r) - \beta \int_{R_2}^\infty \frac{G_{22}(r, r')}{r'} \eta(r') dr' = \frac{\omega}{m} \eta(r). \quad (6.1)$$

Since the kernel  $\frac{G_{22}(r, r')}{r'}$  is symmetric (see Appendix B), the operator on the LHS of the equation acting on  $\eta(r)$  is symmetric; so the eigenvalues are necessarily real. The eigenfunctions are orthogonal with respect to the standard inner product defined by  $\langle f_1, f_2 \rangle = \int_{R_2}^\infty f_1(r) f_2^*(r) dr$  for any two functions  $f_1$  and  $f_2$ , for which this integral is convergent. This integral equation (6.1) is similar in form to the integral equation of the BT flow discussed in detail by ?. Here the Green's function differs because of the cylindrical symmetry and, more importantly, because of the BC component of the Green's function [see (3.18)]. Another difference is the fact that the domain here is unbounded.

Equation (6.1) was solved numerically by using the numerical scheme described at the end of §4.2. The frequency  $\omega$  is indeed always real, and Figure 3 shows some eigenfunction examples. The properties of the spectrum and eigenfunctions are explained analytically below.

Before discussing the structure of the solutions, we make a rough estimate of the allowed frequencies (the spectrum). Multiplication of (6.1) by  $\eta^*$  and integrating yields

$$\int_{R_2}^\infty \frac{\bar{V}_2(r)}{r} |\eta(r)|^2 dr - \beta \int_{R_2}^\infty \int_{R_2}^\infty \frac{G_{22}(r, r')}{r'} \eta^*(r') \eta(r) dr' dr = \frac{\omega}{m} \int_{R_2}^\infty |\eta(r)|^2 dr. \quad (6.2)$$

Figure 3: Examples of perturbation types in the outer region  $r > R_2$ . The shared flow parameters are  $\beta = -0.1$ ,  $\Lambda = 1$ ,  $\lambda_1 = \lambda_2 = 0.5$ ,  $m = 3$ . The arrow designates a delta function. (a) Asymptotically wavelike BT mode with critical layer ( $R_1 = 4$ ,  $R_2 = 5$ ,  $\Gamma_1 = 1$ ), (b) Asymptotically wavelike BT mode without critical layer ( $R_1 = 2$ ,  $R_2 = 3.5$ ,  $\Gamma_1 = -1$ ), (c) Asymptotically evanescent BT mode with critical layer ( $R_1 = 2$ ,  $R_2 = 3.5$ ,  $\Gamma_1 = -1$ ), (d) Asymptotically BC mode without critical layer ( $R_1 = 2$ ,  $R_2 = 3.5$ ,  $\Gamma_1 = -1$ ). Arrows designate delta functions, and their height corresponds to the prefactor multiplying the delta functions.

Since the term  $\frac{G_{22}(r, r')}{r'}$  is always negative [by (3.18), (B 6), and (B 12)], the second integral on the LHS is negative, so the possible values of  $\omega$  are

$$-\infty < \omega < \sup_r \frac{m\bar{V}_2(r)}{r}, \quad (6.3)$$

where “sup” denotes the supremum. For future reference, we define the segments as

$$\mathcal{S}_1 = \left( \inf_r \frac{m\bar{V}_2(r)}{r}, \sup_r \frac{m\bar{V}_2(r)}{r} \right), \quad \mathcal{S}_2 = \left( -\infty, \inf_r \frac{m\bar{V}_2(r)}{r} \right), \quad (6.4)$$

so by (6.3),  $\omega \in \mathcal{S}_1 \cup \mathcal{S}_2$ . We note that, contrary to the bounds derived for the phase velocity given by ? for annular flows (known as the semi-circle theorems, cf. ?), the phase velocity here cannot, by similar arguments, be bounded from below because the flow is unbounded, whereas such theorems use the fact that the flow is confined to a channel (zonal or annular).

### 6.1. Structure of solution near a critical layer

When  $\omega \in \mathcal{S}_1$  there is a critical distance  $r_c$  at which the angular velocity  $\omega/m$  of the perturbation is equal to the angular velocity  $V(r_c)/r_c$  of the flow; the integral equation (6.1) is then singular. In this case, the solution contains a critical layer (see, e.g., Ref. ?). The LHS of the equation can be viewed as the sum of an operator of multiplication by  $\bar{V}_2/r$  and an integral operator. Following Refs. ??, the solution is written in the form of a delta function (the eigenfunction of the multiplication operator) plus an additional term,

$$\eta(r) = D(\omega)\delta\left(\frac{\bar{V}_2}{r} - \frac{\omega}{m}\right) - P \frac{\beta}{\frac{\bar{V}_2}{r} - \frac{\omega}{m}} \xi(r), \quad (6.5)$$

where  $D(\omega)$  and  $\xi(r)$  are unknown functions to be specified,  $\xi(r)$  is assumed to be a regular function of  $r$ ,  $P$  signifies that the principal value of the integral is to be taken when integrating the last expression with respect to  $r$  [i.e.,  $P \int_R^\infty = \lim_{\epsilon \rightarrow 0} (\int_R^{r_c - \epsilon} + \int_{r_c + \epsilon}^\infty)$ ]. Some of the solutions obtained numerically are indeed of the form (6.5), as shown in Figures 3a and 3c; the PV perturbation blows up near the point  $r = r_c$  and a delta function appears at  $r = r_c$ . Note that equation (6.5) is valid also if there is no  $r$  for which  $V(r)/r = \omega/m$ , since then there is no critical layer and we may set  $D(\omega) = 0$ . These regular solutions are shown in Figures 3b and 3d.

Plugging (6.5) into (6.1) yields the following equation for  $\xi$ :

$$\xi(r) = -\frac{D(\omega)G_{22}(r, r_c)}{|(V(r)/r)'|_{r_c}} + P \int_{R_2}^\infty \frac{\beta G_{22}(r, r')}{\left(\frac{\bar{V}_2(r')}{r'} - \frac{\omega}{m}\right) r'} \xi(r') dr', \quad (6.6)$$

where we assume for simplicity that, for  $r > R_2$ , the function  $V_2/r$  is injective and  $r_c = (V_2/r)^{-1}(\omega/m)$ , as is the case for the basic flows considered herein. Also, we used

the mathematical relation  $\delta(f(x)) = \delta(x-x_0)/|f'(x_0)|$  that holds for any smooth injective function  $f(x)$ , where  $x_0$  is a root of  $f(x)$  [if  $x_0$  exists, otherwise  $\delta(f(x)) = 0$ ]. Since (6.1) is homogeneous, we can arbitrarily demand that

$$\int_{R_2}^{\infty} \eta(r') dr' = 1, \quad (6.7)$$

which, by (6.5), is equivalent to the specification of the function  $D(\omega)$  by the following equation:

$$\frac{D(\omega)}{|(V(r)/r)'_{r_c}|} - P \int_{R_2}^{\infty} \frac{\beta}{\frac{\bar{V}_2(r')}{r'} - \frac{\omega}{m}} \xi(r') dr' = 1. \quad (6.8)$$

Using (6.8) in (6.6), we get that  $\xi$  satisfies the following nonsingular inhomogeneous Fredholm equation of the second kind:

$$\xi(r) = -\frac{G_{22}(r, r_c)}{r_c} + \beta \int_{R_2}^{\infty} \frac{G_{22}(r, r')/r' - G_{22}(r, r_c)/r_c}{\frac{\bar{V}_2(r')}{r'} - \frac{\omega}{m}} \xi(r') dr'. \quad (6.9)$$

The nonsingularity is guaranteed since the derivative of the Green's functions is always much less than the derivative of the velocity (inverse power versus linear function of  $r$ ; see Appendixes A and B). Since there is no singularity in this equation at  $r = r_c$ , the function  $\xi(r)$  is regular, as assumed. If  $\omega$  is outside the range of  $\{mV(r)/r\}$ , then the solution consists only of the regular function  $\xi(r)$  with no blowup.

Equation (6.9) can be transformed to a fourth-order nonhomogeneous (homogeneous) differential equation if a critical layer exists (does not exist) by the procedure presented in Appendix D. The nonhomogeneous term (D 6) that appears in the resulting differential equation (D 5) contains only a delta function with its derivatives, which are singular only at  $r = r_c$  (if they exist). Therefore, at  $r \rightarrow \infty$  the solutions to the differential equation asymptotically approach the same form, regardless of whether a critical layer is present. We denote the four linearly independent regular solutions to the equation by  $h_1, h_2, h_3$ , and  $h_4$ . In the following section we find asymptotic expressions for  $h_j$  ( $j = 1, \dots, 4$ ) and find the spectrum properties of the eigenvalue equation (6.1).

## 6.2. Asymptotically barotropic and baroclinic wave types

We now show that there are two types of solution that, asymptotically at large  $r$ , behave as BT and BC waves. For this we resort to the equations in their differential form (3.4) and consider solutions far from the origin, where  $\bar{V}_2$  can be neglected. Because the velocity diminishes exponentially with  $r$  [see (A 20) and (A 21)], such a range may always be found. By (3.4), far from the origin,  $\mathcal{Q}_1 = 0$  and  $\mathcal{Q}_2 = -m\beta\Phi_2/\omega r$ , so equations (3.5) and (2) become

$$0 = \frac{d^2\Phi_1}{dr^2} + \frac{1}{r} \frac{d\Phi_1}{dr} - \frac{m^2}{r^2} \Phi_1 - \frac{\Lambda^2}{\lambda_1} (\Phi_1 - \Phi_2), \quad (6.10)$$

$$-\frac{m\beta\Phi_2}{\omega r} = \frac{d^2\Phi_2}{dr^2} + \frac{1}{r} \frac{d\Phi_2}{dr} - \frac{m^2}{r^2} \Phi_2 + \frac{\Lambda^2}{\lambda_2} (\Phi_1 - \Phi_2). \quad (6.11)$$

Asymptotically we may neglect the LHS of (6.11) since  $\Phi_2$  appears on the RHS without division by  $r$ . We use the ansatz  $\Phi_2 = a\Phi_1$ , where  $a$  is some parameter to be determined. As shown below, this ansatz leads to four independent solutions, which, by the discussion at the end of §6.1, cover all the possible asymptotic solutions of the fourth-order differential equation. Plugging  $\Phi_2 = a\Phi_1$  into (6.10) and (6.12) gives the set of equations

(after dividing the second equation by  $a$ )

$$0 = \frac{d^2\Phi_1}{dr^2} + \frac{1}{r} \frac{d\Phi_1}{dr} - \frac{m^2}{r^2} \Phi_1 - \frac{\Lambda^2}{\lambda_1} (1-a) \Phi_1, \quad (6.12)$$

$$0 = \frac{d^2\Phi_1}{dr^2} + \frac{1}{r} \frac{d\Phi_1}{dr} - \frac{m^2}{r^2} \Phi_1 + \frac{\Lambda^2}{\lambda_2} \frac{(1-a)}{a} \Phi_1. \quad (6.13)$$

These two equations are identical provided that

$$\frac{1-a}{\lambda_2 a} = -\frac{1-a}{\lambda_1} \Rightarrow a = 1 \text{ or } a = -\frac{\lambda_1}{\lambda_2}. \quad (6.14)$$

Thus, asymptotically,  $\Phi_2 \sim \Phi_1$  or  $\Phi_2 \sim -\frac{\lambda_1}{\lambda_2} \Phi_1$ . The first expression corresponds to the asymptotically BT mode, where  $\Phi_{\text{BC}} = \Phi_1 - \Phi_2 \approx 0$ , and the second expression corresponds to the asymptotically BC mode, where  $\Phi_{\text{BT}} = \lambda_1 \Phi_1 + \lambda_2 \Phi_2 \approx 0$ . In the following, we loosely call perturbations whose asymptotic behavior is BT (BC) as BT (BC) modes, without repeating the fact that this behavior is only asymptotic. Also, note that the BT or BC character of the mode is reflected only in the streamfunctions and not in the relations between the PV perturbations, because the PV perturbation at the upper layer is zero in any case; thus, the BT and BC PV perturbations are  $\mathcal{Q}_{\text{BT}} = \lambda_2 \mathcal{Q}_2$  and  $\mathcal{Q}_{\text{BC}} = -\mathcal{Q}_2$  (i.e., they are of the same order of magnitude). Having arrived to the conclusion that there are two kinds of asymptotic modes, we now turn to find their  $r$  dependence.

### 6.2.1. Barotropic mode

First, we assume that the BT component of the streamfunction is the dominant component,  $\Phi_{\text{BT}} \gg \Phi_{\text{BC}}$  (i.e.,  $\Phi_1 \approx \Phi_2$ ). By (3.11), this means that

$$\int_{R_2}^{\infty} \frac{G_{\text{BT}}(r, r')}{r'} \eta(r') dr' \gg \int_{R_2}^{\infty} \frac{G_{\text{BC}}(r, r')}{r'} \eta(r') dr', \quad (6.15)$$

so we take only the BT component of the Green's function in (6.1),

$$\frac{\bar{V}_2(r)}{r} \eta(r) - \beta \int_{R_2}^{\infty} \frac{\lambda_2 G_{\text{BT}}(r, r')}{r'} \eta(r') dr' = \frac{\omega}{m} \eta(r). \quad (6.16)$$

We apply the linear operator  $D_1$  defined by (D 1) on both sides of (6.16) and use (D 2). The term  $D_1(\frac{\bar{V}_2(r)}{r} \eta(r))$  is neglected because the basic velocity and its derivatives are negligible far from the island (see Appendix A). The integral equation is then converted to the differential equation

$$\frac{\eta(r)}{r} = -\frac{\omega}{m\beta\lambda_2} \left( \frac{d^2\eta}{dr^2} + \frac{1}{r} \frac{d\eta}{dr} - \frac{m^2}{r^2} \eta \right). \quad (6.17)$$

If  $\omega < 0$ , the general solution to (6.17) is given by

$$\eta(r) = AH_{2m}^{(1)} \left( 2\sqrt{\frac{m\beta\lambda_2 r}{\omega}} \right) + BH_{2m}^{(2)} \left( 2\sqrt{\frac{m\beta\lambda_2 r}{\omega}} \right), \quad (6.18)$$

where  $H_{2m}^{(1)}$  and  $H_{2m}^{(2)}$  are the Hankel functions of the first and second kinds, respectively, of order  $2m$ . Based on the discussion at the end of §6.2, we denote the two regular solutions to (D 5), for which  $H_{2m}^{(1)}$  and  $H_{2m}^{(2)}$  are their asymptotic approximation, by  $h_1$  and  $h_2$ , respectively. The solutions must obey the radiation condition, according to which energy cannot arrive from outside; this no-radiation condition is satisfied only by  $H_{2m}^{(1)}$ ,

so  $B = 0$  (see ? for details). The solution to (6.16) is then in the form

$$\eta(r) = Ah_1(r; \omega) + D(\omega)\delta(\bar{V}_2/r - \omega/m), \quad (6.19)$$

where asymptotically  $h_1(r; \omega) \sim H_{2m}^{(1)}(2\sqrt{m\beta\lambda_2 r/\omega})$  and  $D(\omega)$  is nonzero if there is a critical layer and zero otherwise. Substitution of (6.19) into (6.16) and applying  $r = R_2$  leads to two options: (i) If  $\omega \in \mathcal{S}_1$  [i.e.,  $D(\omega) \neq 0$ ], then  $A$  is nonzero and is determined by an inhomogeneous equation. Therefore, a solution exists for any  $\omega \in \mathcal{S}_1$ . Such a solution, having a critical layer and asymptotically BT, is shown in Figure 3a. (ii) Conversely, if no critical layer exists, then  $D(\omega) = 0$ , and the equation is homogeneous in  $A$ . Therefore, in this case,  $\omega$  takes only discrete values in segment  $\mathcal{S}_2$ . Such an asymptotically BT solution without a critical layer is shown in Figure 3b.

If  $\omega > 0$ , the general solution to (6.17) is given by a superposition of the modified Bessel functions of order  $2m$ ,

$$\eta(r) = \tilde{A}K_{2m}\left(2\sqrt{-\frac{m\beta\lambda_2 r}{\omega}}\right) + \tilde{B}I_{2m}\left(2\sqrt{-\frac{m\beta\lambda_2 r}{\omega}}\right). \quad (6.20)$$

These two functions are the asymptotic approximations to  $h_1$  and  $h_2$  in this case. For the solutions to be limited as  $r \rightarrow \infty$ , we must set  $\tilde{B} = 0$ . In virtue of (6.3), the case  $\omega > 0$  occurs only if  $\bar{V}_2$  is positive, in which case it approaches zero at infinity (see Appendix A). Therefore,  $\omega \in \mathcal{S}_1$  and this type of perturbation always contains a critical layer:

$$\eta(r) = Ah_1(r; \omega) + D(\omega)\delta(\bar{V}_2/r - \omega/m), \quad (6.21)$$

with  $D(\omega) \neq 0$ . An example of this solution is shown in Figure 3c. Substitution of (6.21) into (6.16) and using  $r = R_2$  leads to  $A \neq 0$  with no limitation on  $\omega$ . Therefore,  $\omega$  can take any value in the segment  $\mathcal{S}_1$ .

### 6.2.2. Baroclinic mode

We now consider the case where the BC component dominates. In this case, by (6.1),

$$\int_{R_2}^{\infty} \frac{\lambda_1 G_{\text{BC}}(r, r')}{r'} \eta(r') dr' = -\frac{\omega}{m\beta} \eta(r). \quad (6.22)$$

We apply the linear operator  $D_2$  defined by (D1) on both sides of (6.22) and, by using (D2), obtain

$$\frac{\eta(r)}{r} = -\frac{\omega}{m\beta\lambda_1} \left( \frac{d^2 \eta}{dr^2} + \frac{1}{r} \frac{d\eta}{dr} - \frac{m^2}{r^2} \eta - \frac{\Lambda^2}{\lambda_1 \lambda_2} \eta \right). \quad (6.23)$$

The general solution to (6.23) is

$$\eta(r) = \frac{E}{\sqrt{r}} W_{\kappa, m} \left( \frac{2\Lambda r}{\sqrt{\lambda_1 \lambda_2}} \right) + \frac{F}{\sqrt{r}} M_{\kappa, m} \left( \frac{2\Lambda r}{\sqrt{\lambda_1 \lambda_2}} \right), \quad (6.24)$$

where  $\kappa = \frac{m\beta\lambda_1\sqrt{\lambda_1\lambda_2}}{2\Lambda\omega}$ ,  $W_{\kappa, m}$  and  $M_{\kappa, m}$  are the Whittaker functions of order  $(\kappa, m)$ , and  $E$  and  $F$  are constants. By using the asymptotic form of the Whittaker functions (?), we get asymptotically

$$\eta(r) \sim Er^{\kappa - \frac{1}{2}} e^{-\frac{\Lambda r}{\lambda_1 \lambda_2}} + Fr^{-\kappa - \frac{1}{2}} \frac{\Gamma(1 + 2m)}{\Gamma(\frac{1}{2} + m - \kappa)} e^{\frac{\Lambda r}{\lambda_1 \lambda_2}}. \quad (6.25)$$

For the solution to be limited as  $r \rightarrow \infty$ , we set  $F = 0$ . Note that  $W_{\kappa, m}$  has positive zeros only if  $\kappa > 1/2$  (?). This inequality yields  $m\beta\sqrt{\lambda_1\lambda_2}/\Lambda \leq \omega < 0$ , i.e., only negative

Segment containing $\omega$	Sign of $\omega$	Asymptotic mode	Asymptotic form	Spectrum is continuous or discrete
$\mathcal{S}_1$	$\omega > 0$	BT	$CK_{2m}(2\sqrt{m\beta\lambda_2 r/\omega})$	continuous
$\mathcal{S}_1$	$\omega < 0$	BT	$AH_{2m}^{(1)}(2\sqrt{m\beta\lambda_2 r/\omega})$	continuous
$\mathcal{S}_1$	any $\omega \neq 0$	BC	$EW_{\kappa,m}(2\Lambda r/\sqrt{\lambda_1\lambda_2})/\sqrt{r}$	continuous
$\mathcal{S}_2$	$\omega < 0$	BT	$AH_{2m}^{(1)}(2\sqrt{m\beta\lambda_2 r/\omega})$	discrete
$\mathcal{S}_2$	any $\omega \neq 0$	BC	$EW_{\kappa,m}(2\Lambda r/\sqrt{\lambda_1\lambda_2})/\sqrt{r}$	discrete

Table 1: Parts of the spectrum and the associated properties. The segments  $\mathcal{S}_1$  and  $\mathcal{S}_2$  are defined by (6.4). BT and BC designate BT mode and BC mode, respectively.

eigenvalues yield wavelike asymptotic eigenfunctions, an example of which is shown in Figure 3d.

Referring now to the discussion at the end of §6.2, we denote by  $h_3$  and  $h_4$ , respectively, the two regular solutions to (D 5), with  $W_{\kappa,m}(2\Lambda r/\sqrt{\lambda_1\lambda_2})/\sqrt{r}$  and  $M_{\kappa,m}(2\Lambda r/\sqrt{\lambda_1\lambda_2})/\sqrt{r}$  being their asymptotic approximation. The general solution to (6.22) is then of the form

$$\eta(r) = Eh_3(r; \omega) + D(\omega)\delta(\bar{V}_2/r - \omega/m), \quad (6.26)$$

where asymptotically  $h_3(r; \omega) \sim H_{2m}^{(1)}(2\sqrt{m\beta\lambda_2 r/\omega})$  and  $D(\omega)$  is nonzero if there is a critical layer and zero otherwise. Substitution of (6.26) into (6.22) and applying  $r = R_2$  leads, as in the BT case, to two options: (i) If  $\omega \in \mathcal{S}_1$  [i.e.,  $D(\omega) \neq 0$ ], then  $E$  is nonzero and is determined by an inhomogeneous equation. Therefore, a solution exists for any  $\omega \in \mathcal{S}_1$ . (ii) Conversely, if no critical layer exists, then  $D(\omega) = 0$ , and the equation is homogeneous in  $E$ . Therefore, in this case,  $\omega$  can take only discrete values in the segment  $\mathcal{S}_2$ . Such an asymptotically BC solution without a critical layer is shown in Figure 3d. The solution is wavelike in some region and then, starting from some distance, decays exponentially in  $r$ , as implied by (6.25).

A summary of the different parts of the spectrum is listed in Table 1. One part consists of all the values in segment  $\mathcal{S}_1$  (excluding zero), where each value has multiplicity 2 (i.e., there are two corresponding eigenfunctions with a critical layer). These eigenfunctions correspond asymptotically to BT or BC forms. These are evanescent if  $\omega > 0$ . If  $\omega < 0$ , the asymptotically BT type is wavelike as  $r \rightarrow \infty$ , and the asymptotically BC type is wavelike in a finite region if  $\omega > m\beta\sqrt{\lambda_1\lambda_2}/\Lambda$  and otherwise is evanescent. The rest of the spectrum is a discrete set of the segment  $(-\infty, \inf\{m\bar{V}_2/r\})$ , including asymptotically BT and BC types without a critical layer.

### 6.3. Decay of asymptotically BT and BC modes

The modal analysis above in §6.2 shows that the perturbation types belonging to the continuous spectrum are neutral (i.e., are maintained without growth or damping with time). However, a correct treatment of the initial-value problem correctly shows that such modes may give rise to asymptotic algebraic decay with time (?) or to algebraic growth (see, e.g., Ref. ?). Since the perturbation expressions in the complex- $\omega$  plane contain poles and branch cuts, as seen from (6.18) or (6.24), a natural question is what is their contribution to the flow stability properties? As shown in this section, these types contribute to its stability by causing decay rather than neutrality of a given initial perturbation.

Consider a time-dependent PV perturbation of azimuthal mode number  $m$  in the lower

Figure 4: (a) Schematic drawing of the location on the complex- $\omega$  plane of the poles and branch lines of the Laplace transform response. There are poles due to the discrete spectrum [where  $D(\omega) = 0$ ], a pole at  $\omega = m\bar{V}_2(r_0)/r_0$ , a pole at  $\omega = m\bar{V}_2(r)/r$ , a pole at  $\omega = 0$  due to the asymptotically BC mode, a branch line at  $\mathcal{S}_1$ , and a branch line at  $\text{Im}(\omega) < 0$  due to the asymptotically BT mode. Also, the Bromwich contour  $\text{Im}(\omega) = \gamma$  is designated. (b) Contour for calculating the inverse Laplace transform of the asymptotically BT mode.

layer,  $q_2(r, \theta, t) = \zeta_2(r, t)e^{im\theta}$ . Its Laplace transform is

$$\mathcal{Q}_2(r, \omega) = \int_0^\infty \zeta_2(r, t)e^{i\omega t} dt, \quad (6.27)$$

where the notation  $\mathcal{Q}_i$  is consistent with the definition (3.3). The inverse Laplace transform is

$$\zeta_2(r, t) = \frac{1}{2\pi} \int_{-\infty+i\gamma}^{\infty+i\gamma} \mathcal{Q}_2(r, \omega)e^{-i\omega t} d\omega, \quad (6.28)$$

where the Bromwich contour of integration is along  $\text{Im}(\omega) = \gamma$ , where  $\gamma$  is greater than the imaginary part of all the singularities of  $\mathcal{Q}_2(r, \omega)$ .

Laplace-transforming the linearized equation for  $q_2$  in (3.2) gives

$$\left(\frac{\bar{V}_2}{r} - \frac{\omega}{m}\right) \mathcal{Q}_2 - \frac{\Phi_2}{r} \frac{d\bar{Q}_2}{dr} = \frac{\zeta_2(r, t=0)}{im}. \quad (6.29)$$

Assume for simplicity that  $\zeta_2(r, t=0) = \delta(r - r_0)/r$  and denote the solution to (6.29) in this case by  $\chi(r; r_0; \omega)/r$ . This solution is the response function of the system to an initial delta-function perturbation. Thus, the equation for  $\chi(r; r_0, \omega)$  is

$$\left(\frac{\bar{V}_2}{r} - \frac{\omega}{m}\right) \chi(r; r_0, \omega) - \beta \int_{R_2}^\infty \frac{G_{22}(r, r')}{r'} \chi(r'; r_0, \omega) dr' = \delta(r - r_0). \quad (6.30)$$

Asymptotically, as  $r \rightarrow \infty$  the solutions to (6.30) coincide with the solutions to (6.1), which by §6.2 asymptotically take the form of Hankel functions of the first kind [see (6.18)] or Whittaker functions [see (6.24)]. Since, for large  $\omega$ ,

$$H_{2m}^{(1)}\left(2\sqrt{\frac{m\beta r}{\omega}}\right) \sim \frac{1}{(2m)!} \left(\frac{m\beta r}{\omega}\right)^{2m}, \quad W_{\kappa, m}\left(\frac{2\Lambda r}{\sqrt{\lambda_1 \lambda_2}}\right) \sim r^{\frac{m\beta \lambda_1 \sqrt{\lambda_1 \lambda_2}}{2\Lambda \omega} - \frac{1}{2}} e^{-\frac{\Lambda r}{\lambda_1 \lambda_2}}, \quad (6.31)$$

then  $\chi$  is bounded as  $|\omega| \rightarrow \infty$ . It is therefore possible to deform the Bromwich contour integral until it consists only of integrals around poles and cuts.

Appendix E shows that the poles of  $\chi(r; r_0; \omega)/r$  are of four types: (i) a discrete isolated set corresponding to perturbation types with no critical layer, (ii) the point  $\omega = \bar{V}_2(r_0)/r_0$ , (iii) a branch cut along segment  $\mathcal{S}_1$ , and (iv) the poles of the regular functions  $\xi(r; r_0, \omega)$  defined by (E3). Here we use the asymptotic ( $r \rightarrow \infty$ ) expressions for  $\xi(r; r_0, \omega)$ , which are identical to the asymptotic perturbations found in §6.2. Two types of singularities occur in the asymptotic regime: one is the singularity  $1/\sqrt{\omega}$  that appears in (6.18), and the second is the singularity  $1/\omega$  that appears in (6.24) (in the expression for  $\kappa$ ). To account for the singularity  $1/\sqrt{\omega}$ , a branch of the square root must be chosen; for convenience we choose the branch cut to be on the negative imaginary axis. A schematic drawing illustrating the various poles and branch-cut locations is given



in Figure 4a. The singularity at  $\omega = 0$  and the branch cut of  $\sqrt{\omega}$  are unique to the beta-cone model.

The contribution of (i) to the inverse Laplace transform is a discrete sum of exponentials of the form  $e^{-i\omega_n t}$ , where  $\{\omega_n\}$  is the discrete set mentioned. In the same way, the pole at  $m\omega = \bar{V}_2(r_0)/r_0$  gives rise to a simple exponential  $e^{-im\bar{V}_2(r_0)t/r_0}$ . The branch cut  $\mathcal{S}_1$  results in algebraic decay as  $1/t$  (?).

To calculate the contribution of the BT mode to the integral, we use the integration shown in Figure 4b. The integral along the small circle,  $\int_0^{2\pi} H_{2m}^{(1)}(2\sqrt{m\beta r/\epsilon}e^{i\phi})e^{-i\epsilon e^{i\phi}t}d\phi$ , vanishes as  $\epsilon \rightarrow 0$  (this can be found by direct numerical integration). Denoting the negative imaginary axis by the frequency  $\omega = ix$ , where  $x$  is real, the contribution to the integral (6.28) along the right side of the branch cut is

$$\int_{-\infty}^0 H_{2m}^{(1)}\left(2\sqrt{\frac{m\beta r}{ix}}\right)e^{-iixt}dx = -\int_0^{\infty} H_{2m}^{(1)}\left(2\sqrt{\frac{m\beta r}{-ix}}\right)e^{-xt}dx. \quad (6.32)$$

Asymptotically as  $t \rightarrow \infty$ , significant contributions to the integral come only from points  $x$  near zero. Therefore, the Hankel function in the integrand can be replaced by its asymptotic approximation at  $x \sim 0$ , which is

$$H_{2m}^{(1)}\left(2\sqrt{\frac{m\beta r}{-ix}}\right) \sim \left(\frac{-i}{m\pi^2\beta r}\right)^{1/4} x^{1/4} \exp\left(2i\sqrt{\frac{m\beta r}{-ix}} - im\pi - \frac{i\pi}{4}\right). \quad (6.33)$$

(?). Since the exponential term is bounded by 1, the integral in (6.32) is bounded by the following integral:

$$\left(\frac{1}{m\pi\beta r}\right)^{1/4} \int_0^{\infty} x^{1/4}e^{-xt}dx = \left(\frac{1}{m\pi\beta r}\right)^{1/4} t^{-5/4}\Gamma(5/4), \quad (6.34)$$

where  $\Gamma$  is the gamma function. The integral over the other line gives an identical time dependence, so we conclude that the perturbation decays asymptotically as  $t^{-5/4}$  in this case.

The contribution of the asymptotically BC mode is simpler because there is only one singularity at  $\omega = 0$  with no branch cuts. We assume that  $r$  is large enough so the asymptotic expansion of the Whittaker function can be used,  $W_{\kappa,m}(r) \sim r^{\frac{\Omega}{\omega} - \frac{1}{2}}e^{-\frac{\Lambda r}{\lambda_1 \lambda_2}}$  (?), where  $\Omega = \frac{m\beta\lambda_1\sqrt{\lambda_1\lambda_2}}{2\Lambda}$ . The inverse Laplace transform is then

$$\frac{1}{2\pi} \int_{-\infty+i\gamma}^{\infty+i\gamma} W_{\kappa,m}(r)d\omega \propto \frac{1}{2\pi} \int_{-\infty+i\gamma}^{\infty+i\gamma} r^{\frac{\Omega}{\omega}} e^{-i\omega t} d\omega = i\delta(t) + \frac{i\sqrt{i\Omega}}{\sqrt{t}} J_1\left(2\sqrt{|\Omega|t \ln r}\right), \quad (6.35)$$

where the last equality is from Ref. ?. For large times,  $J_1(2\sqrt{|\Omega|t \ln r}) \sim t^{-1/4} \cos(2\sqrt{|\Omega|t \ln r} - 3\pi/4)$  (?), so the BC mode oscillates while its amplitude decays as  $t^{-3/4}$ .

## 7. Aspects of contour-contour and contour-topography instabilities

### 7.1. Contour-contour instability

In the CC resonance, the instability is due to the interaction of the PV waves at the liquid contours  $r = R_1$  and  $r = R_2$ . In this case, the bottom topography at  $r > R_2$  can be neglected; this amounts to setting  $\beta = 0$  where it appears explicitly in (4.9)–(4.11) [but not setting  $\beta = 0$  in the expressions for PV discontinuities  $\Delta_1$  and  $\Delta_2$  in (4.2)]. By (4.11),  $\eta$  vanishes in this case; a system of two homogeneous algebraic equations for  $\alpha_1$

and  $\alpha_2$  is established. This system can be written in matrix form,

$$\begin{bmatrix} M_{11} - \frac{\omega}{m} & M_{12} \\ M_{21} & M_{22} - \frac{\omega}{m} \end{bmatrix} \begin{bmatrix} \alpha_1 \\ \alpha_2 \end{bmatrix} = \begin{bmatrix} 0 \\ 0 \end{bmatrix}, \quad (7.1)$$

where

$$M_{11} = \frac{\bar{V}_1(R_1) - \Delta_1 G_{11}(R_1, R_1)}{R_1}, \quad M_{12} = -\frac{\Delta_1 G_{12}(R_1, R_2)}{R_2}, \quad (7.2)$$

$$M_{21} = -\frac{\Delta_2 G_{21}(R_2, R_1)}{R_1}, \quad M_{22} = \frac{\bar{V}_2(R_2) - \Delta_2 G_{22}(R_2, R_2)}{R_2}. \quad (7.3)$$

To have a nontrivial solution, the determinant of the  $2 \times 2$  matrix in (7.1) should be zero. This yields the eigenvalue equation, which is quadratic in  $\omega$ :

$$\frac{\omega^2}{m^2} - (M_{11} + M_{22})\frac{\omega}{m} + M_{11}M_{22} - M_{12}M_{21} = 0, \quad (7.4)$$

from which we get the dispersion relation for the two-contour subsystem:

$$\omega_{A,B} = \frac{m}{2} \left[ (M_{11} + M_{22}) \pm \sqrt{(M_{11} + M_{22})^2 - 4(M_{11}M_{22} - M_{12}M_{21})} \right]. \quad (7.5)$$

The subscript A or B corresponds to applying a + or – sign before the square root in (7.5), respectively. The two eigenvectors corresponding to the two eigenvalues in (7.5) are the two modes of PV perturbations at the liquid contours, which we call type A or type B; they are connected to the contour deformations via (4.6) (see Refs. ??).

Figure 5 shows an example of how the CC instability can be recognized via the dispersion curves  $\omega(m)$ . The basic flow parameters are  $R_1 = R_2 = 2.5$ ,  $\Gamma_1 = -1$ ,  $\lambda_1 = \lambda_2 = 1/2$ ; for future reference, this setup is called “configuration A” of the flow. The eigenvalues of the isolated CC system are calculated by using (7.5), while the eigenvalues of the full system are calculated numerically as explained in §4.2. To facilitate tracking of the dispersion relation, we calculate the dispersion curves for continuously varying  $m$  and mark the points corresponding to integer  $m$ , which are the physically relevant values [see (3.3)]. Figure 5a shows the angular phase velocity of the perturbations,  $\text{Im}(\omega)/m$  versus the wave number  $m$ , for the two CC waves and the unstable perturbation. When  $m = 3$ , 4, or 5, the two phase velocities of the two CC waves coincide and CC instability occurs, as can be seen from the curve for growth rate  $[\text{Im}(\omega)]$  in Figure 5b.

At  $m = 6$ , the angular velocities of the two CC waves differ, so no CC instability is possible. However, because the flow remains unstable at  $m = 6$  (the growth rate is nonzero), we conclude that one of the CC waves is in resonance with the topographic perturbations at  $r > R_2$  (because the outer region at  $r > R_2$  by itself is always stable, see §6).

For a CC instability, the growth rate of the full system  $g_F$  can be compared with that found by using the isolated CC system  $g_{CC}$ . For the case shown in Figure 5b, at low mode number ( $m = 3$  or  $m = 4$ ), the inequality  $g_F \leq g_{CC}$  holds; therefore, at these mode numbers, the topography outside causes a reduction in the growth rate. At  $m = 6$  the inequality is reversed,  $g_F > g_{CC} = 0$ , the topography at  $r > R_2$  then destabilizes the flow, since without it the flow would stay stable. This result is not specific to the particular parameters of the flow in this example but occurred in all our calculations: at low mode numbers the CC resonance is dominant, yet the full-system growth rate is lower than expected because of the CC interaction alone. Also, at larger mode numbers the CT resonance becomes the only one that contributes to instability while the CC subsystem is stable. Another example of this result is given below.

Figure 5: Real and imaginary parts of the eigenvalues for CC and full resonance. The basic flow parameters are  $R_1 = 2.5$ ,  $R_2 = 2.5$ ,  $\Gamma_1 = -1c$ ,  $\beta = -0.5$  (configuration A). (a) Perturbation angular velocity  $\text{Re}(\omega)$  versus the mode number  $m$ . The angular velocities of two CC waves are given by red and blue dotted curves, and that of the full system is given by the green curve whenever there is instability. Points with physically relevant values of  $m$  (integers) are marked. S, CC, and CT designate regions where the flow is stable, unstable because of CC resonance, and unstable because of CT resonance, respectively. (b) Growth rate  $\text{Im}(\omega)$  versus the mode number for the CC resonance ( $g_{CC}$ , purple) and for the full system ( $g_F$ , green). The points are joined by straight lines for better visualization.

### 7.2. Contour-topography instability

As shown in Figure 5, the unstable  $m = 6$  mode, which is not caused by a CC resonance, has a real angular velocity very close to that of one of the CC-interaction modes. This suggests that the CC perturbation type B, having the lowest angular velocity of the two modes [see (7.5)], is actually the mode that resonates with the topographic perturbations for  $r > R_2$ . To identify the resonating perturbation type in the CC system by the integral eigenvalue approach, we rewrite the eigenvalue equations (4.9)–(4.11) so that the CC perturbation types appear decoupled; the calculation is given in Appendix F. This procedure can be viewed as partial diagonalization of the system of equations (4.9)–(4.10) by moving to the CC eigenmode coordinates. The resulting equations (G 5)–(G 7) in Appendix F are diagonal in the isolated CC system (i.e., in case there is no topography outside the contours). The results indicate that, for the CCT instability, only type B is in resonances with the topographic perturbations.

To understand why type B is the mode that resonates with the topographic perturbations, we apply pseudomomentum considerations. Recall that two modes may resonate only if their pseudomomenta are of opposite sign (§5.1). Because the topographic types have positive pseudomomentum [see (5.4)], only the type having negative pseudomomentum can resonate with them. Appendix F proves that the pseudomomentum of a perturbation has the same sign as the slope of the dispersion curve (when  $m$  may be taken to vary smoothly); a similar result was obtained for shallow, rotating water (one layer, zonal) in Ref. ?. Figure 5a shows clearly that, at  $m = 6$ , only type B has negative pseudomomentum because its dispersion curve is the only one that decreases with  $m$  at  $m = 6$ .

As shown in Figure 2b, a flow with parameters  $R_1 = 5$ ,  $R_2 = 2$ ,  $\Gamma_1 = -1$ ,  $\lambda_1 = \lambda_2 = 0.5$  (we call this flow configuration B) is unstable against mode  $m = 2$  perturbations, where the instability is the CT instability. Figure 6 shows that  $m = 2$  is the gravest unstable mode in this case. Again, the full-system phase velocity is close to that of perturbation type B of the CC subsystem, which is consistent with its decreasing dispersion curve, pointing the fact that it has negative pseudomomentum (Appendix F).

Figure 7 shows the growth rates of different mode numbers as functions of the radius  $R_2$  of the lower ring for the basic flow parameters  $R_1 = 5$ ,  $\lambda_1 = \lambda_2 = 0.5$ , and  $\beta = -0.1$ . When  $\Gamma_1 = 1$  (Figure 7a), the lines of the  $m \geq 2$  modes are composed of two “bulges” that get close to each other with increasing mode number until they merge at  $m = 8$ . The instability in this bulge regime is of type CC, as shown for the  $m = 2$  case in Figure 2a. On the left of each of the lines, it becomes nearly horizontal; in this range the instability is of type C<sub>2</sub>T (this is also shown for  $m = 2$  in Figure 2a). Between these two regions the

Figure 6: Real and imaginary parts of the eigenvalues for CC and full resonance. The basic flow parameters are  $R_1 = 5$ ,  $R_2 = 2$ ,  $\Gamma_1 = -1$ ,  $\beta = -0.1$  (configuration B). (a) Perturbation angular velocity  $\text{Re}(\omega)$  versus mode number  $m$ . (b) Growth rate  $\text{Im}(\omega)$  versus mode number for the full system. The notation and colors are the same as in Figure 5.

Figure 7: Growth rates  $\text{Im}(\omega)$  for different mode numbers as functions of  $R_2$  at  $R_1 = 5$ ,  $\beta = -0.1$ ,  $\Lambda = 1$ , and  $\lambda_1 = \lambda_2 = 0.5$  for (a)  $\Gamma_1 = 1$ , (b)  $\Gamma_1 = -1$ . Each curve is labeled by its mode number.

instability is of type CCT. Unlike the other modes, mode  $m = 1$  is unstable only because of the CC resonance.

Similarly, when  $\Gamma = -1$  (Figure 7b), the lines of the growth rates for  $m \geq 2$  modes are composed of three parts: one is the low- $R_2$  regime, where the lines are nearly horizontal, in which case the instability is of type  $C_1T$ . The CC instability part consists of the line where a steep increase in growth rate begins (going from left to right). Between these two regions the instability is of type CCT; this was also shown for the  $m = 2$  case in Figure 2b. Again, mode  $m = 1$  is unstable only because of CC resonance.

### 7.3. Barotropic and baroclinic contour-topography resonance

Another useful property of the eigenvalue equations in integral form, (4.9)–(4.11), is the simple separation of BT and BC couplings. By equations (3.15)–(3.18), the Green's functions  $G_{11}$ ,  $G_{12}$ ,  $G_{21}$ , and  $G_{22}$  are linear combinations of the two more basic BC and BT Green functions,  $G_{BT}$  and  $G_{BC}$ . The latter serve as coupling coefficients between the contours' perturbations  $\alpha_1$  and  $\alpha_2$  and the perturbation  $\eta$  outside. Therefore, if we use  $G_{BT} \equiv 0$  ( $G_{BC} \equiv 0$ ) in the integral terms in (4.9)–(4.11), only BC (BT) couplings to the outside perturbation are allowed. Upon comparing the resulting growth rates for any case we can identify which of the couplings is dominant. When the BT (BC) coupling is dominant, the contours enter in resonance with the asymptotically BT (BC) mode. If both couplings are dominant, the contours enter in resonance with a mixed mode.

In most cases, the BT CT resonance dominates, whereas the BC CT resonance is very weak or absent. Figure 8a shows an example of the growth rates of the full system, the CC subsystem, the BT coupling, and the BC coupling; the relative thickness  $\lambda_1$  of the upper layer is varied. In this case,  $R_1 = 5$ ,  $R_2 = 2$ ,  $\Gamma_1 = -1$ ,  $\beta = -0.1$ , and the mode number is  $m = 2$ , so the instability is of type CCT (see Figure 2b) and the type-B perturbation of the CC subsystem enters resonance with the BT perturbation type whenever an instability occurs.

An example of a configuration where the BC coupling dominates is not easily found because, for all  $r$  and  $r'$ , the inequality  $G_{BC}(r, r') < G_{BT}(r, r')$  holds, as may be verified directly from (B 6) and (B 12). Moreover, if  $|r - r'| \gg r$ , this inequality gets stronger:  $G_{BC}(r, r') \ll G_{BT}(r, r')$ . Therefore, the BC interaction terms are usually negligible compared with the BT terms. As a result, the growth rates and the eigenfunctions are determined mainly by the BT couplings. This explains the BT governor effect, in which BT shear reduces the BC growth rate (?). In the case of circularly symmetric flow, ? noted that this effect may be attributed to BT strain rather than to shear, but both the shear  $\partial\bar{V}_2/\partial r$  and strain  $r\partial(\bar{V}_2/r)/\partial r$  are nonzero in the present case.

The only example where the BC CT resonance dominates is when the BT growth rates approach zero. In this case, the subsystem with only BT couplings can be viewed

as almost stable, and the BC couplings can be viewed as small perturbations that can affect the resulting eigenvalues of the full system. Figure 8b shows such an example, where  $R_1 = R_2 = 5$ ,  $\Gamma_1 = 1$ , and  $\beta = -0.1$ . In the range  $0.123 \leq \lambda_1 \leq 0.15$  the instability is of type C<sub>2</sub>T, in which case the growth rate of the full system  $g_F$  is very close to that due to the BC coupling  $g_{BC}$ .

As  $\lambda_1$  approaches unity (i.e., the lower layer becomes very thin) the BC coupling becomes dominant (Figure 8b). This is consistent with the findings of Ref. ? who investigated two-layer shallow water with a bottom topography and found that, as the depth of the lower layer decreases, the BC instability overtakes the BT instability. In terms of the resonance perturbations, however, this range of  $\lambda_1 \approx 1$  cannot be attributed to instability with BC perturbations because the instability is CC in this range (because  $g_{CC} \neq 0$ ); the BC interaction only contributes to increasing the growth rate, not to its origin.

#### 7.4. Resonance with continuous spectrum

Figure 9 shows examples of three unstable solutions (found numerically) for  $\eta$  for three configurations of the basic flow. Figure 9a shows the basic flow profiles for configuration A (see §7.1 and Figure 5), and Figures 9d and 9g show the amplitude and relative phase of the resulting unstable perturbation, respectively. Because  $\omega_r = \text{Re}(\omega) < 0$  in this case while  $\bar{V}_2 \geq 0$ , this cannot be a critical layer instability. The instability is of type CC and the growing perturbation outside is reminiscent of the asymptotically BT mode (6.20). Given that  $\omega$  is complex, the alternating PV profile decreases exponentially with  $r$  (for details see Ref. ?).

Figure 9b shows the PV and velocity profiles of configuration B (see §7.1 and Figure 6). As shown above, the instability is of type CCT for this configuration, and because  $\omega_r \in \mathcal{S}_1$ , this is a critical layer instability. The unstable perturbation (Figure 9e) is reminiscent of the critical layer structure (§4), as is most clearly viewed by the rapid change in the relative phase of the perturbations over a thin region near  $r_c = m(\bar{V}_2/r)^{-1}(\omega_r)$  (Figure 9h).

The third configuration, which we label configuration C, consists of a C<sub>2</sub>T instability and a dominant BC coupling. The flow parameters are  $R_1 = R_2 = 5$ ,  $\Gamma_1 = 1$ ,  $\beta = -0.1$ ,  $\lambda_1 = 0.14$ , and  $\lambda_2 = 0.86$ . Figure 8 shows clearly the dominance of the BC coupling. The unstable perturbations (Figures 9f and 9i) are reminiscent of the critical layer structure near  $r_c$  (at about 5.1) and, at  $r > r_c$ , it recalls the stable BC mode.

To understand the structure of the solutions for CT instability we approximate the solution to the eigenvalue equations when the growth rates are small. The integral equation (4.11) is nonsingular and  $\eta(r)$  is then given by (6.5) with no delta function and without having to calculate the principal value:

$$\eta(r) = \frac{\xi(r)}{\frac{\bar{V}_2}{r} - \frac{\omega}{m}}. \quad (7.6)$$

Plugging (7.6) into (4.11) yields the following equation for  $\xi$ :

$$-\frac{G_{21}(r, R_1)}{R_1} \alpha_1 - \frac{G_{22}(r, R_2)}{R_2} \alpha_2 + \xi(r) = \int_{R_2}^{\infty} \frac{\beta G_{22}(r, r')}{\frac{\bar{V}_2(r')}{r} - \frac{\omega}{m}} \xi(r') dr'. \quad (7.7)$$

We now use  $\omega = \omega_r + i\omega_i$ , where  $\omega_r$  and  $\omega_i$  are the real and imaginary parts of  $\omega$ , respectively. If  $\omega$  is near a bifurcation (i.e.,  $\omega_i$  is small), then we may assume that  $\text{Im} \xi$  is

Figure 8: Growth rates  $\text{Im}(\omega)$  for different resonances as functions of  $\lambda_1$  for (a)  $R_1 = 5$ ,  $R_2 = 2$ ,  $\Gamma_1 = -1$ ,  $\beta = -0.1$ ,  $m = 2$ ; (b)  $R_1 = R_2 = 5$ ,  $\Gamma_1 = 1$ ,  $\beta = -0.1$ ,  $m = 2$ . Each curve is labeled by resonance type. The inset in the upper left of panel (b) shows the growth rates in the range  $0.12 < \lambda_1 < 0.16$ .

Figure 9: Examples of (a)–(c) profiles of the basic flow, (d)–(f) the corresponding unstable PV perturbations amplitude, and (g)–(i) the relative phases. The basic flow parameters for each vertical column are (a), (d), (g)  $R_1 = R_2 = 2.5$ ,  $\Gamma_1 = -1$ ,  $\beta = -0.5$ ,  $\lambda_1 = \lambda_2 = 0.5$  (configuration A), where  $m = 5$  is the gravest mode with frequency  $\omega = -0.118 + 0.081i$ . (b), (e), (h)  $R_1 = 5$ ,  $R_2 = 2$ ,  $\Gamma_1 = -1$ ,  $\beta = -0.1$ ,  $\lambda_1 = \lambda_2 = 0.5$  (configuration B), where  $m = 2$  is the gravest mode with frequency  $\omega = 0.221 + 0.027i$ . (c), (f), (i)  $R_1 = R_2 = 5$ ,  $\Gamma_1 = 1$ ,  $\beta = -0.1$ ,  $\lambda_1 = 0.2$ ,  $\lambda_2 = 0.8$  (configuration C), where  $m = 2$  is the gravest mode with frequency  $\omega = -0.0097 + 0.0002i$ . Notations and colors of Figures (a)–(c) are the same as in Figure 1. In panels (d)–(f) the PV perturbations in the upper layer are denoted by solid blue lines and in the lower layer by dotted blue lines. Arrows denote delta functions, their height corresponds to the prefactors of the delta functions.

also small. By (3.3) the expression for the PV perturbation at  $r > R_2$  is

$$q_2(r, \theta, t) = -\frac{\beta \text{Re}[\xi(r)] \left( \frac{\bar{V}_2}{r} - \frac{\omega_r}{m} \right)}{r \left( \frac{\bar{V}_2}{r} - \frac{\omega_r}{m} \right)^2 + \frac{\omega_i^2}{m^2}} e^{\omega_i t} \cos(m\theta - \omega t), \quad (7.8)$$

where the term  $\text{Im}[\xi(r)]\omega_i$  was neglected because it is second order in  $\omega_i$ . The solution vanishes at  $r = m\bar{V}_2/\omega_r$  and changes the sign of PV between the two sides; note that the similarity to a critical layer structure is more prominent as the ratio  $\omega_i/\omega_r$  becomes small.

For the basic flow in Figures 9b and 9c, the ratio is  $\omega_i/\omega_r \approx 0.128$  and  $\omega_i/\omega_r \approx 0.1$ , respectively. By (7.8), as  $\omega_i \rightarrow 0$ , the region over which the PV changes sign in the unstable mode becomes more narrow; thus, in the limit of  $\omega_i \rightarrow 0$ , the discontinuous nature of the critical layer is restored. In the approximation of linear perturbations, equation (7.8) describes a thin layer having an  $m$ -fold symmetry centered at  $r = m\bar{V}_2/\omega_r$  and that becomes stronger and broader with time.

In the case of CT resonance, some of the eigenmodes in the outer region  $r > R_2$  form a continuum (§6). This suggests that the resonance in this case is with a collection of perturbations of the continuous spectra, as was shown in Ref. ? and as may be explained simply based on pseudomomentum considerations: the pseudomomentum of the resonating perturbations in the system must sum to zero (see §5.1). Since the pseudomomentum of the contours is always finite [the first two terms in (5.4)], the pseudomomentum of the topographic perturbation outside must also be finite. However, the pseudomomentum of one critical-layer perturbation [i.e., the third term in (5.4)] is infinite by (6.5), so any finite-pseudomomentum topographic perturbation must be composed of a collection of critical-layer perturbations such that the third term in (5.4) is finite.

Following Ref. ?, we use projections to determine the structure of this collection. The unstable outer PV perturbation in the lower layer,  $\eta_\omega(r)$  [where  $\text{Im}(\omega) > 0$ ], is projected on the possible stable self-excitations of the outer region discussed in §4,  $\eta_{\omega'}$  {where  $\omega'$  is real and the critical layer is at  $r_c = (V_2/r)^{-1}(\omega'/m)$ }. Because the stable solutions constitute an orthonormal set (see §4.1), the projection  $\langle \eta_\omega, \eta_{\omega'} \rangle$  correctly calculates the

weights in this collection. By using (6.5) we get

$$\begin{aligned}
 \langle \eta_\omega, \eta_{\omega'} \rangle &= \int_{R_2}^{\infty} \eta_\omega(r) \left[ D(\omega') \delta \left( \frac{\bar{V}_2(r)}{r} - \frac{\omega'}{m} \right) + P \frac{\xi_{\omega'}^*(r)}{\bar{V}_2 - \frac{\omega'}{m}} \right] dr \\
 &= \frac{D(\omega')}{\left| (\bar{V}_2/r)'_{r_c} \right|} \eta_\omega(r_c(\omega')) + P \int_{R_2}^{\infty} \frac{\xi_\omega(r) \xi_{\omega'}^*(r) dr}{\left( \frac{\bar{V}_2}{r} - \frac{\omega}{m} \right) \left( \frac{\bar{V}_2}{r} - \frac{\omega'}{m} \right)} \\
 &\approx \frac{D(\omega')}{\left| (\bar{V}_2/r)'_{r_c(\omega')} \right|} \eta_\omega(r_c(\omega')) - \frac{i\pi}{\left| (\bar{V}_2/r)'_{r_c(\omega')} \right|} \eta_\omega(r_c(\omega')) \xi_{\omega'}^*(r_c(\omega')).
 \end{aligned} \tag{7.9}$$

The principal value integral was calculated by using Cauchy's integral theorem (see, e.g., Ref. ?) and the fact that  $\omega_i > 0$  for an unstable mode, and by assuming that the main contribution to the integral is near the critical layer. We assume that  $D(\omega')$ ,  $\xi_{\omega'}(r_c(\omega'))$ , and  $\bar{V}_2(r_c(\omega'))$  depend weakly on  $\omega'$  relative to  $\eta_\omega(r_c(\omega'))$ . In this case, the weight goes as

$$|\langle \eta_\omega, \eta_{\omega'} \rangle|^{1/2} \sim |\eta_\omega(r_c(\omega'))|^{1/2} = \left[ \frac{|\xi_\omega(r_c(\omega'))|}{(\omega' - \omega_r)^2 + \omega_i^2} \right]^{1/2}. \tag{7.10}$$

If  $\xi_\omega(r_c(\omega'))$  depends weakly enough on  $\omega'$ , then (7.10) is maximized at  $\omega' = \omega_r$ , which means that the resonance eigenfunctions are those with frequencies close to  $\omega_r$ . Figure 10 shows that the approximation (7.10) is consistent with the direct calculation of the weight. Approximating (7.10) further by assuming that the numerator is constant leads to

$$|\langle \eta_\omega, \eta_{\omega'} \rangle|^{1/2} \sim \left[ \frac{1}{(\omega' - \omega_r)^2 + \omega_i^2} \right]^{1/2}, \tag{7.11}$$

as obtained by Ref. ? in the single-layer, zonal, rotating shallow-water model. However, this expression is not a good approximation for frequencies outside the immediate neighborhood of  $\omega_r$ , as shown in Figure 10.

## 8. Nonlinear evolution of contour-contour versus contour-topography instabilities

According to the linear stability analysis scheme, it seems at first glance that no substantial difference should exist between the evolution of the flow for a CC instability and that for a CT instability. After all, the source of the instability causes the entire system to collectively rotate and grow; the unstable solutions of the linear stability analysis are written as if phase-locking occurs immediately. However, in practice, phase-locking evolves over time (see, e.g., Ref. ?). If the system is subject to some random noise, the first two parts to phase-lock are the resonance perturbation types. Therefore, over time, they grow into the dominant perturbations, and the remaining perturbations are influenced by these initial perturbations. The subsystem in resonance is the first to have large-scale perturbations, so nonlinear effects become pronounced first for this subsystem.

We now discuss high-Reynolds-number simulations done with the coefficient-form partial differential equation package of COMSOL, which is based on the finite-element method (see Ref. ? for details). The vorticity-diffusion term  $\nu \nabla^2 Q_i$  is added to the RHS of (2.5) to maintain numerical stability. The resulting coupled system composed of (2.4) and the equation for PV evolution [i.e., (2.5) supplemented with the diffusion term] is solved as an initial-value problem in a two-dimensional  $(r, \theta)$  rectangular grid with the limits  $1 < r < 30$  and  $0 \leq \theta \leq 2\pi$ . The unknown variables are the streamfunction and the

Figure 10: Amplitude of spectrum of PV perturbation in lower layer at  $r > R_2$  (blue solid line), approximate expression (7.10) (red dotted line), and approximate expression (7.11) (green dotted line) for  $R_1 = 5$ ,  $R_2 = 2$ ,  $\Gamma_1 = -1$ ,  $\beta = -0.1$  (configuration B).

Figure 11: Evolution of relative PV field in upper layer (upper panel in each pair) and lower layer (lower panel in each pair). The basic flow parameters are  $R_1 = R_2 = 2.5$ ,  $\Gamma_1 = -1$ ,  $\beta = -0.1$ , and  $\lambda_1 = \lambda_2 = 0.5$  (configuration A); mode 5 is the most unstable. Red and blue designate positive and negative PV, and gray designates the island. Time is specified in dimensionless units at the upper-left corner of the upper panel in each pair.

PVs. We apply the periodicity conditions at  $\theta = 0$  and  $\theta = 2\pi$  and the no-slip conditions at both radial boundaries by setting  $\partial\Psi_i/\partial r = \partial\Psi_i/\partial\theta = 0$  at  $r = 30$ , and  $\partial\Psi_i/\partial r = 0$  and  $\Psi_i = 0$  at  $r = 1$ .

The computational domain is  $30 \times 2\pi$  in size and is divided into three subdomains. The first is a fine-grid domain that covers  $1 \leq r < 1.5$  with a mesh size of  $0.05 \times 0.03$ ; it is set off to resolve the viscous boundary layer that may form next to the cylinder. The second is the main domain and covers  $1.5 \leq r < 20$  with a mesh size of  $0.1 \times 0.03$ . In both these domains,  $\nu$  is set to 0.0001. The third domain covers  $20 \leq r \leq 30$  and is set off as an absorbing layer to prevent reflections. To obtain a reasonable machine time for the evolution of linear instability of the flow, we add over the entire computational grid a random perturbation to the basic PV field in the form of Gaussian noise. More details on the method used are available in Ref. ?.

### 8.1. Contour-contour instability

Figure 11 shows an example of the evolution of an unstable flow for a CC instability. The flow parameters are the same as for Figures 5 and 9a; namely,  $R_1 = 2.5$ ,  $R_2 = 2.5$ ,  $\Gamma_1 = -1$ , and  $\beta = -0.5$ . As shown in Figure 5b, the gravest unstable mode is  $m = 5$ , which is indeed the mode that evolves most rapidly in the simulation. Initially, the contours deform at  $t = 30$ ; the upper contour tilts relative to the lower contour, and they are phase-locked and propagate clockwise in accordance with the calculated frequency of the linear stability analysis,  $\omega = -0.118 + 0.081i$ .

As shown at  $t = 60$  and  $t = 80$ , the contour perturbations excite the perturbation types outside in the form of waves, which are similar to the stationary waves discussed in §6.2. Since the BC wave mode decreases exponentially with distance, the BT wave dominates, which decreases only as  $r^{-1/4}$ . This BT mode has the form of spirals, as was shown for BT flows on the beta cone (?). These linear waves appear only in the lower layer, where the gradient of the basic PV exists.

During the nonlinear growth of the deformation, five pairs of partially overlapping negative (in the upper layer) and positive (in the lower layer) PV patches can be identified (most clearly at  $t = 60$ ). These can be viewed as modons, i.e., QG BC vortical dipoles (see, e.g., Ref. ???). For each pair, the positive part remains attached to the cylinder, whereas the negative part is released and moves more freely (as can be seen for time  $t \geq 160$ ). Initially ( $t = 90$ ), each pair is created such that it propagates toward the cylinder, which causes the positive patch to deform and approach the cylinder. The positive patch then changes partner and the dipole moves outwards again ( $t = 120$ ). Upon reaching a maximal distance from the island, the modons swing ( $t = 160$ ) and return to the cylinder. Next, the modons collide, exchange partners, and new modons emerge. Because of wave radiation, dissipation, and filamentation, the maximal distance



from the cylinder is smaller this time. This process repeats in a quasi-periodic manner in a fashion similar to the BT evolution shown in Ref. ?. At  $t = 230$  the fivefold symmetry is lost, two of the positive parts in the upper layer leave the cylinder, and the five modons are wandering around.

The evolution in this case of CC instability is very similar to that of unstable BT flows on the beta cone studied by ?. The main features of the QG evolution are the emergence of modons (instead of dipoles in the BT case), which tend to move counterclockwise; the appearance of spiral BT PV waves propagating clockwise; and a quasiperiodic outward and inward motion of modons that exchange partners every cycle. The ‘‘average’’ beta in this system can be defined according to the weight of each layer as  $\lambda_1 \cdot 0 + \lambda_2 \cdot \beta = \lambda_2 \beta$ , which gives  $-0.25$  for the evolution in Figure 11. As expected and as found in simulations (data not shown), the maximal distance of the modons increases for weaker  $|\beta|$ , whereas, for stronger  $|\beta|$ , new flow patterns form without the emergence of modons.

### 8.2. Contour-topography instability

Whereas the CC instability produces flow evolutions analogous to those in the BT case, the CT instability is rather different. The main reason is that the dominant interaction is now between one of the contours and a perturbation at  $r > R_2$ . This perturbation recalls the critical layer solution discussed in §7.4, so the contour is in resonance with a thin layer of PV with alternating sign located some distance from it. The parts in resonance are those with the greatest PV-perturbation amplitude and thus are the first to reach nonlinear saturation during the phase-locking stage (cf. Ref. ?).

Figure 12 shows the evolution of CCT instabilities. The flow is in configuration B, as for Figures 6 and 9b, where  $R_1 = 5$ ,  $R_2 = 2$ ,  $\Gamma_1 = -1$ , and  $\beta = -0.1$ . The resonance is between type-B perturbations of the CC subsystem and the topographic perturbations for  $r > R_2$ . As shown in Figure 7b, the CCT instability in this configuration is close to the  $C_1$ T-instability regime, so the mutual deformation of the contours significantly deforms  $C_1$  ( $t = 30$ ), whereas  $C_2$  suffers only minor deformations. By  $t = 30$ , a narrow PV ring with  $m = 2$  symmetry forms at  $r \approx 4.2$ ; this ring is the collection of critical-layer perturbations with critical layers in the vicinity of  $r = 4.2$  (§7.4).

Out of the initial random perturbations that were inserted into the system, only the resonant perturbations begin to phase-lock and grow. Therefore, the  $C_1$  deformation and the new thin PV ring are the first to grow significantly in this case and reach large-scale perturbations, where nonlinearity becomes important. Nonlinear effects stop the linear growth and the thin ring rearranges in the configuration shown at  $t = 90$ . This cessation of growth explains why no dipolar modons emerge, contrary to the CC-instability case (§8.1), where both contours are significantly deformed.

From  $t = 90$  to  $t = 230$ , the flow rotates counterclockwise and completes about three revolutions in a quasi-stationary manner. This structure is a BC version of the tripolar structure found in the BT beta-cone model (?, Figure 14) and recalls the stationary two-layer QG tripole vortices found numerically on the  $f$  plane (??) and that were also investigated on the beta plane (?).

The tripole eventually breaks into two modon quartets ( $t = 260$ ) composed of two PV patches at each layer. In the upper layer, the PVs of the circle and its adjacent patch are equal, approximately  $-1$ , as was initially the case. In the lower layer, the PVs differ because only the PV of the circular patch was there initially, its value being approximately  $8.32$ . The PV of the adjacent patch, which emerges from the interaction with the upper-layer PV, is  $0.3$  on average. Because the positive-PV circular core in the lower layer is so strong and only the noncircular patch in the upper layer is tilted vertically relative to it, this quartet behaves effectively as a dipolar modon (composed

Figure 12: Evolution of relative PV field in the upper layer (upper panel in each pair) and lower layer (lower panel in each pair). The basic flow parameters are  $R_1 = 5$ ,  $R_2 = 2$ ,  $\Gamma_1 = -1$ ,  $\beta = -0.1$ , and  $\lambda_1 = \lambda_2 = 0.5$  (configuration B); mode 2 is the most unstable. Colors and notation are the same as in Figure 11.

of a circular positive vortex in the lower layer and a noncircular vortex patch in the upper layer). Therefore, after reaching a maximal distance of about 2.8 from the island ( $t = 260$ ), the modon swings and returns to the cylinder ( $t = 294$ ).

## 9. Conclusion

We investigate herein the possible resonances leading to instability of two-layer QG circular flows around an island with the sea bottom sloping offshore ( $\beta < 0$ ). The flow in each layer is composed of one uniform relative PV ring: the outer radius of the upper (lower) ring is  $R_1$  ( $R_2$ ) and the dimensionless PV inside it is  $\Gamma_1 = +1$  or  $-1$  [ $\Gamma_2$ , given by (4.3)]. An azimuthal normal-mode analysis leads to a set of integral eigenvalue equations that have direct physical interpretation in terms of the possible resonances of the system.

Topographic PV perturbations are possible only in the lower layer at  $r > R_2$ , where a nonzero PV gradient occurs. A continuous set of possible perturbations consists of those having a critical layer. Asymptotically, as  $r \rightarrow \infty$ , these solutions split into two types: BT or BC modes. When these modes rotate clockwise, they are wavelike in the radial direction, so a pattern of spiral PV patches appear on the two-dimensional plane. Although both modes seem to be neutral in the normal analysis scheme, the full initial-value treatment shows that they actually decay over time.

At low mode numbers (usually  $m = 2, 3, 4$  for the cases discussed herein), the CC resonance dominates the CT interactions, yet the full-system growth rate is less than expected because of the CC interaction alone. Thus, coupling to the external topographic perturbations stabilizes the system. At larger mode numbers, the CT interaction becomes the only unstable interaction.

For a fixed radius  $R_1$  of the upper layer ring, the radius  $R_2$  of the lower layer ring determines the type of resonance that leads to instability. When the lower ring is sufficiently thin (i.e.,  $R_2$  approaches 1), the dominant resonance is CT. If  $\Gamma_1 = +1$  (in which case the flow in both layers is clockwise), then the resonance is specifically  $C_2T$ , which means that the lower ring contour is in resonance with the topographic perturbations outside of it. If  $\Gamma_1 = -1$  (in which case the flow in both layers is counterclockwise), then the resonance is specifically  $C_1T$ , which means that the upper ring contour is in resonance with the topographic perturbations at  $r > R_2$ . The transition from small  $R_2$ , where the instability is  $C_1T$  (or  $C_2T$ ), to large  $R_2$ , where the instability is CC, occurs through the CCT instability. In this instability, one of the perturbation types of the CC subsystem enters resonance with the topographic perturbations, but not  $C_1$  (or  $C_2$ ) itself.

The resonance between the contours and the topographic perturbations may be dominated either by BT or BC couplings, which are easy to identify by using the integral-equation approach. Usually, the BT couplings dominate (the BT governor effect), but for a narrow region of the upper layer of relative thickness  $\lambda_1$ , the dominant instability is due to baroclinity. In this case, the resonance of the contours is primarily with the asymptotic BC topographic mode.

The nature of the instability reflects the stage of nonlinear evolution of the flow. In case of CC instability, the two contours change significantly during the phase-locking stage, which leads to modon formation and emission from the island. In the CT instability

the resonance involves a collection of topographic perturbations with critical layers in proximity to one another. The result is a strengthening of the PV in a narrow ring at some distance in the lower layer. This ring interacts with the contours to form a quasistationary structure (e.g., a tripole) and only later breaks into modons that may be emitted from the island.

With some minor modifications, the beta-cone concept can be used to treat flows in the presence of the conical beta effect on a planetary scale, namely, for the Antarctic Circumpolar Current. In this case, (2.2) for the upper-layer PV is supplemented with an additional background planetary beta term  $\beta_P r$  (related to the gradient of the Coriolis parameter), and  $\beta$  in the lower layer is replaced by  $\beta_P + \beta_T$  ( $\beta_T$  being related to the bottom topography); see, e.g., Ref. ?. In this case more resonances come into play, because more types of perturbations are added in the upper layer. These issues will be considered separately elsewhere.

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## Appendix A. Velocity profile of basic flow

Consider the basic flow, in which the PV in each of the layers is given by (4.1). The equations can be decoupled by using the following definitions of the BT and BC modes of the basic flow (cf. ?):

$$\bar{Q}_{BT} = \lambda_1 \bar{Q}_1 + \lambda_2 \bar{Q}_2, \quad \bar{\Psi}_{BT} = \lambda_1 \bar{\Psi}_1 + \lambda_2 \bar{\Psi}_2, \quad \beta_{BT} = \lambda_2 \beta, \quad (\text{A } 1)$$

$$\bar{Q}_{BC} = \bar{Q}_1 - \bar{Q}_2, \quad \bar{\Psi}_{BC} = \bar{\Psi}_1 - \bar{\Psi}_2, \quad \beta_{BC} = -\beta. \quad (\text{A } 2)$$

From (A 1) and (A 2) we obtain

$$\bar{Q}_1 = \bar{Q}_{BT} + \lambda_2 \bar{Q}_{BC}, \quad \bar{\Psi}_1 = \bar{\Psi}_{BT} + \lambda_2 \bar{\Psi}_{BC}, \quad (\text{A } 3)$$

$$\bar{Q}_2 = \bar{Q}_{BT} - \lambda_1 \bar{Q}_{BC}, \quad \bar{\Psi}_2 = \bar{\Psi}_{BT} - \lambda_1 \bar{\Psi}_{BC}, \quad \beta_2 = \beta_{BT} - \lambda_1 \beta_{BC}. \quad (\text{A } 4)$$

By using (A 3) and (A 4) along with (2.4), we arrive at the equations that relate the modal PVs and streamfunctions:

$$\bar{Q}_{BT} = \nabla^2 \bar{\Psi}_{BT} + \beta_{BT} r, \quad (\text{A } 5)$$

$$\bar{Q}_{BC} = \nabla^2 \bar{\Psi}_{BC} - \tilde{\Lambda}^2 \bar{\Psi}_{BC} + \beta_{BC} r, \quad (\text{A } 6)$$

where  $\tilde{\Lambda} = \Lambda / \sqrt{\lambda_1 \lambda_2}$ . For definiteness, we assume that  $R_2 > R_1$ ; otherwise the following expressions should be adapted in a straightforward manner. By using (A 5) and (4.1), the BT streamfunction satisfies the equation

$$\bar{\Psi}_{BT}'' + \frac{1}{r} \bar{\Psi}_{BT}' + \beta_{BT} r = \begin{cases} \lambda_1 \Gamma_1 + \lambda_2 \Gamma_2, & R \leq r \leq R_1 \\ \lambda_2 \Gamma_2, & R_1 < r \leq R_2 \\ \beta_{BT} r, & R_2 < r. \end{cases} \quad (\text{A } 7)$$

The general solution to (A 7) is

$$\bar{\Psi}_{BT} = \begin{cases} -\frac{1}{9} \beta_{BT} r^3 + \frac{1}{4} (\lambda_1 \Gamma_1 + \lambda_2 \Gamma_2) r^2 + C_1 \ln(r) + C_2, & R \leq r \leq R_1 \\ -\frac{1}{9} \beta_{BT} r^3 + \frac{1}{4} \lambda_2 \Gamma_2 r^2 + C_3 \ln(r) + C_4, & R_1 < r \leq R_2 \\ C_5 + C_6 \ln r, & R_2 < r. \end{cases} \quad (\text{A } 8)$$

The expression for the energy of the flow is (see, e.g., Ref. ?)

$$E = \frac{1}{2} \iint_{r>R} [\lambda_1(\nabla\bar{\Psi}_1)^2 + \lambda_2(\nabla\bar{\Psi}_2)^2] r dr d\theta + \frac{1}{2} \Lambda^2 \iint_{r>R} (\Psi_1 - \Psi_2)^2 r dr d\theta, \quad (\text{A } 9)$$

where the first integral gives the kinetic energy and the second integral gives the available potential energy of the flow. The potential energy is associated with the BC mode only, whereas the kinetic energy is contributed by both modes. Therefore, for the kinetic energy to be finite, the contribution to the first integral in (A 9) from the BT mode should be finite and we must set  $C_6 = 0$  in (A 8). By (A 7) the BT streamfunction is continuous, as is its first derivative:

$$\bar{\Psi}_{\text{BT}}(R_1^-) = \bar{\Psi}_{\text{BT}}(R_1^+), \quad \bar{\Psi}_{\text{BT}}(R_2^-) = \bar{\Psi}_{\text{BT}}(R_2^+), \quad (\text{A } 10)$$

$$\bar{\Psi}'_{\text{BT}}(R_1^+) = \bar{\Psi}'_{\text{BT}}(R_1^-), \quad \bar{\Psi}'_{\text{BT}}(R_2^+) = \bar{\Psi}'_{\text{BT}}(R_2^-). \quad (\text{A } 11)$$

From equations (A 10) and (A 11) the four unknowns  $C_1$ – $C_4$  are

$$C_1 = (1/3)\lambda_2\beta R_2^3 - (1/2)\Gamma_1 R_1^2 \lambda_1 - (1/2)\lambda_2\Gamma_2 R_2^2, \quad (\text{A } 12)$$

$$C_2 = (1/2)\ln(R_1)R_1^2\Gamma_1\lambda_1 - (1/4)R_1^2\Gamma_1\lambda_1 - (1/3)\ln(R_2)\lambda_2\beta R_2^3 + (1/2)\ln(R_2)\lambda_2\Gamma_2 R_2^2 + (1/9)\lambda_2\beta R_2^3 - (1/4)\lambda_2\Gamma_2 R_2^2 + C_5, \quad (\text{A } 13)$$

$$C_3 = (1/3)\lambda_2\beta R_2^3 - (1/2)\lambda_2\Gamma_2 R_2^2, \quad (\text{A } 14)$$

$$C_4 = -(1/3)\ln(R_2)\lambda_2 * \beta R_2^3 + (1/2)\ln(R_2)\lambda_2\Gamma_2 R_2^2 + (1/9)\lambda_2\beta R_2^3 - (1/4)\lambda_2\Gamma_2 R_2^2 + C_5. \quad (\text{A } 15)$$

Using (A 1), (A 2), and (A 8), the azimuthal BT velocity is

$$\bar{V}_{\text{BT}} \equiv \frac{\partial \bar{\Psi}_{\text{BT}}}{\partial r} = \begin{cases} -\frac{1}{3}\beta_{\text{BT}}r^2 + \frac{1}{2}(\lambda_1\Gamma_1 + \lambda_2\Gamma_2)r + \frac{C_1}{r}, & R \leq r \leq R_1 \\ -\frac{1}{3}\lambda_2\beta_2r^2 + \frac{1}{2}\lambda_2\Gamma_2r + \frac{C_3}{r}, & R_1 < r \leq R_2 \\ 0, & R_2 < r. \end{cases} \quad (\text{A } 16)$$

The velocity is assumed to vanish at  $r = R$ ; i.e.,  $\bar{V}_{\text{BT}}(R) = 0$  (see §4.1). Applying (A 16) imposes the relation between  $\Gamma_1$  and  $\Gamma_2$  that appears in (4.3). Based on the Stokes theorem this is equivalent to the condition of vanishing total BT excess PV (i.e., the PV that results from omitting the background PV) in the two rings:  $\int_R^{R_2} r \nabla^2 \bar{\Psi}_{\text{BT}} dr = \lambda_1 \int_R^{R_1} r \Gamma_1 dr + \lambda_2 \int_R^{R_2} r (\Gamma_2 - \beta r) dr = 0$ .

The use of (A 6) and (4.1) shows that the BC streamfunction satisfies the equation

$$\bar{\Psi}''_{\text{BC}} + \frac{1}{r}\bar{\Psi}'_{\text{BC}} - \tilde{\Lambda}^2\bar{\Psi}_{\text{BC}} + \beta_{\text{BC}}r = \begin{cases} \Gamma_1 - \Gamma_2, & R \leq r \leq R_1 \\ \beta_1r - \Gamma_2, & R_1 < r \leq R_2 \\ \beta_{\text{BC}}r, & R_2 < r. \end{cases} \quad (\text{A } 17)$$

The general solution to (A 17) is

$$\bar{\Psi}_{\text{BC}} = \begin{cases} D_1 K_0(\tilde{\Lambda}r) + D_2 I_0(\tilde{\Lambda}r) - (\Gamma_1 - \Gamma_2)/\tilde{\Lambda}^2 + i s_{2,0}(i\tilde{\Lambda}r)\beta_{\text{BC}}/\tilde{\Lambda}^3, & R \leq r \leq R_1 \\ D_3 K_0(\tilde{\Lambda}r) + D_4 I_0(\tilde{\Lambda}r) + \Gamma_2/\tilde{\Lambda}^2 - i s_{2,0}(i\tilde{\Lambda}r)\beta_2/\tilde{\Lambda}^3, & R_1 < r \leq R_2 \\ D_5 K_0(\tilde{\Lambda}r) + D_6 I_0(\tilde{\Lambda}r), & R_2 < r, \end{cases} \quad (\text{A } 18)$$

where  $s_{2,0}$  is the Lommel function  $s$  of order  $\{2, 0\}$  (?). For the energy (A 9) to be finite, we must set  $D_6 = 0$ . The BT streamfunction satisfies the continuity conditions at  $r = R_1$  and  $r = R_2$ , the continuity conditions of its derivative at these radii [the equations corresponding to (A 10) and (A 11) for the BC mode], and the condition that

it must vanish at  $r = R$ . By solving these five equations and using the relations

$$\frac{d}{dr}K_0(\tilde{\Lambda}r) = -\tilde{\Lambda}K_1(\tilde{\Lambda}r), \quad \frac{d}{dr}I_0(\tilde{\Lambda}r) = \tilde{\Lambda}I_1(\tilde{\Lambda}r), \quad \frac{d}{dr}s_{2,0}(i\tilde{\Lambda}r) = -\frac{1}{2}i\tilde{\Lambda}\pi L_1(\tilde{\Lambda}r) \quad (\text{A } 19)$$

( $L_1$  being the modified Struve function (?)), the expressions for  $D_1$ – $D_5$  can be found (not given here). The azimuthal BC velocity is then

$$\bar{V}_{\text{BC}} \equiv \frac{\partial \bar{\Psi}_{\text{BC}}}{\partial r} = \begin{cases} -D_1\tilde{\Lambda}K_1(\tilde{\Lambda}r) + D_2\tilde{\Lambda}I_1(\tilde{\Lambda}r) + \pi L_1(\tilde{\Lambda}r)\beta_{\text{BC}}/\tilde{\Lambda}^2, & R \leq r \leq R_1 \\ -D_3\tilde{\Lambda}K_1(\tilde{\Lambda}r) + D_4\tilde{\Lambda}I_1(\tilde{\Lambda}r) - \pi L_1(\tilde{\Lambda}r)\beta_2/\tilde{\Lambda}^2, & R_1 < r \leq R_2 \\ -D_5\tilde{\Lambda}K_1(\tilde{\Lambda}r), & R_2 < r. \end{cases} \quad (\text{A } 20)$$

The basic velocity in each layer is then found by using

$$V_1 = V_{\text{BT}} + \lambda_2 V_{\text{BC}}, \quad V_2 = V_{\text{BT}} - \lambda_1 V_{\text{BC}}, \quad (\text{A } 21)$$

which follow from the definitions above (A 1), (A 2), (A 16), and (A 20).

## Appendix B. Barotropic and baroclinic Green's functions

The BT Green's function  $G_{\text{BT}}(r, r')$  is defined by

$$\frac{d^2 G_{\text{BT}}(r, r')}{dr^2} + \frac{1}{r} \frac{dG_{\text{BT}}(r, r')}{dr} - \frac{m^2}{r^2} G_{\text{BT}}(r, r') = \delta(r - r') \quad (\text{B } 1)$$

and satisfies the boundary conditions

$$G_{\text{BT}}(r = R, r') = 0, \quad G_{\text{BT}}(r \rightarrow \infty, r') = 0. \quad (\text{B } 2)$$

The general solution to (B 1) is

$$G_{\text{BT}} = \begin{cases} ar^m + br^{-m}, & R \leq r < r' \\ cr^m + dr^{-m}, & r' < r. \end{cases} \quad (\text{B } 3)$$

Imposing the boundary conditions (B 2) we get  $b = -aR^{2m}$  and  $c = 0$ . By (B 1) the Green's function is continuous at  $r = r'$ ,

$$G_{\text{BT}}(r'^+, r') = G_{\text{BT}}(r'^-, r'). \quad (\text{B } 4)$$

Integration of (B 1) in the neighborhoods of the singularity  $r = r'$  yields

$$G_{\text{BT}}(r'^+, r') - G_{\text{BT}}(r'^-, r') = 1. \quad (\text{B } 5)$$

Using (B 4) and (B 5), the coefficients  $a$ ,  $b$ , and  $d$  in (B 3) are found. The solution is

$$G_{\text{BT}}(r, r') = \begin{cases} \frac{r'^{-m+1}(R^{2m}r^{-m} - r^m)}{2m}, & R \leq r \leq r' \\ \frac{r'^{-m+1}(R^{2m} - r'^{2m})}{2m} r^{-m}, & r' < r. \end{cases} \quad (\text{B } 6)$$

In the same manner, the BC Green's function  $G_{\text{BC}}$  is defined by the equation

$$\frac{d^2 G_{\text{BC}}(r, r')}{dr^2} + \frac{1}{r} \frac{dG_{\text{BC}}(r, r')}{dr} - \frac{m^2}{r^2} G_{\text{BC}}(r, r') - \frac{\Lambda^2}{\lambda_1 \lambda_2} G_{\text{BC}}(r, r') = \delta(r - r'), \quad (\text{B } 7)$$

and satisfies the boundary conditions

$$G_{\text{BC}}(r = R, r') = 0, \quad G_{\text{BC}}(r \rightarrow \infty, r') = 0. \quad (\text{B } 8)$$

The general solution to (B 7) is (denoting  $\tilde{\Lambda} = \Lambda/\sqrt{\lambda_1\lambda_2}$ )

$$G_{\text{BC}}^m = \begin{cases} \tilde{a}K_m(\tilde{\Lambda}r) + \tilde{b}I_m(\tilde{\Lambda}r), & R \leq r \leq r' \\ \tilde{c}K_m(\tilde{\Lambda}r) + \tilde{d}I_m(\tilde{\Lambda}r), & r' < r. \end{cases} \quad (\text{B } 9)$$

Imposing the boundary conditions (B 8), we get  $\tilde{b} = -\tilde{a}K_m(\tilde{\Lambda}R)/I_m(\tilde{\Lambda}R)$  and  $\tilde{d} = 0$ . By (B 7) the Green's function is continuous at  $r = r'$ ,

$$G_{\text{BC}}(r'^+, r') = G_{\text{BC}}(r'^-, r'). \quad (\text{B } 10)$$

Integration of (B 7) in the neighborhoods of the singularity  $r = r'$  yields

$$G_{\text{BC}}(r'^+, r') - G_{\text{BC}}(r'^-, r') = 1. \quad (\text{B } 11)$$

Using (B 10) and (B 11) and the identity  $I_m K_{m+1} + I_{m+1} K_m = 1/r$  (?), we get the solution

$$G_{\text{BC}}(r, r') = \begin{cases} r'[I_m(\tilde{\Lambda}R)K_m(\tilde{\Lambda}r) - I_m(\tilde{\Lambda}r)K_m(\tilde{\Lambda}R)] \frac{K_m(\tilde{\Lambda}r')}{\tilde{\Lambda}K_m(\tilde{\Lambda}R)}, & R \leq r \leq r' \\ r'[I_m(\tilde{\Lambda}R)K_m(\tilde{\Lambda}r') - I_m(\tilde{\Lambda}r')K_m(\tilde{\Lambda}R)] \frac{K_m(\tilde{\Lambda}r)}{\tilde{\Lambda}K_m(\tilde{\Lambda}R)}, & r' < r. \end{cases} \quad (\text{B } 12)$$

We note that both  $\frac{G_{\text{BT}}(r, r')}{r'}$  and  $\frac{G_{\text{BC}}(r, r')}{r}$  are symmetric with respect to switching of the variables  $r$  and  $r'$ ; this fact is used in §6.

### Appendix C. Pseudomomentum continuity equation

Substituting (4.6) into (3.2) gives

$$\frac{\partial s_1}{\partial t} + \frac{\bar{V}_1}{r} \frac{\partial s_1}{\partial \theta} + \frac{1}{r} \frac{\partial \psi_1}{\partial \theta} = 0, \quad \frac{\partial s_2}{\partial t} + \frac{\bar{V}_2}{r} \frac{\partial s_2}{\partial \theta} + \frac{1}{r} \frac{\partial \psi_2}{\partial \theta} = 0. \quad (\text{C } 1)$$

Multiplying both equations (C 1) by  $s_i \frac{dQ_i}{dr}$  and integrating azimuthally gives

$$\frac{1}{2} \frac{d\bar{Q}_1}{dr} \frac{\partial}{\partial t} \int_0^{2\pi} s_1^2 d\theta + \frac{1}{2} \frac{d\bar{Q}_1}{dr} \frac{\bar{V}_1}{r} \int_0^{2\pi} \frac{\partial s_1^2}{\partial \theta} d\theta - \frac{1}{r} \int_0^{2\pi} q_1 \frac{\partial \psi_1}{\partial \theta} d\theta = 0, \quad (\text{C } 2)$$

$$\frac{1}{2} \frac{d\bar{Q}_2}{dr} \frac{\partial}{\partial t} \int_0^{2\pi} s_2^2 d\theta + \frac{1}{2} \frac{d\bar{Q}_2}{dr} \frac{\bar{V}_2}{r} \int_0^{2\pi} \frac{\partial s_2^2}{\partial \theta} d\theta - \frac{1}{r} \int_0^{2\pi} q_2 \frac{\partial \psi_2}{\partial \theta} d\theta = 0. \quad (\text{C } 3)$$

The second integrals in (C 2) and (C 3) vanish identically. Multiplying (C 2) by  $\lambda_1$ , (C 3) by  $\lambda_2$ , and summing gives

$$\frac{1}{2} \frac{\partial}{\partial t} \int_0^{2\pi} \left( \lambda_1 r \frac{dQ_1}{dr} s_1^2 + \lambda_2 r \frac{dQ_2}{dr} s_2^2 \right) d\theta = \int_0^{2\pi} \left( \lambda_1 q_1 \frac{\partial \psi_1}{\partial \theta} + \lambda_2 q_2 \frac{\partial \psi_2}{\partial \theta} \right) d\theta. \quad (\text{C } 4)$$

Since  $q_1 = \nabla^2 \psi_1 - \frac{\Lambda^2}{\lambda_1} (\psi_1 - \psi_2)$  and  $q_2 = \nabla^2 \psi_2 + \frac{\Lambda^2}{\lambda_2} (\psi_1 - \psi_2)$  [by (2.4)] the RHS of (C 4) takes the form (some of the integrals vanish identically),

$$\int_0^{2\pi} \left( \lambda_1 \nabla^2 \psi_1 \frac{\partial \psi_1}{\partial \theta} + \lambda_2 \nabla^2 \psi_2 \frac{\partial \psi_2}{\partial \theta} \right) d\theta. \quad (\text{C } 5)$$

The first term in the integral may be written as

$$\nabla^2 \psi_1 \frac{\partial \psi_1}{\partial \theta} = \frac{1}{r} \frac{\partial}{\partial r} \left( r \frac{\partial \psi_1}{\partial \theta} \frac{\partial \psi_1}{\partial r} \right) - \frac{1}{2} \frac{\partial}{\partial \theta} \left( \frac{\partial \psi_1}{\partial r} \frac{\partial \psi_1}{\partial r} \right) + \frac{1}{2r^2} \frac{\partial}{\partial \theta} \left( \frac{\partial \psi_1}{\partial \theta} \frac{\partial \psi_1}{\partial \theta} \right), \quad (\text{C } 6)$$

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and upon substituting (C 6) into (C 4), we get

$$-\frac{1}{2} \frac{\partial}{\partial t} \int_0^{2\pi} \left( \lambda_1 r \frac{dQ_1}{dr} s_1^2 + \lambda_2 r \frac{dQ_2}{dr} s_2^2 \right) d\theta + \frac{1}{r} \frac{\partial}{\partial r} \int_0^{2\pi} \left( \lambda_1 \frac{\partial \psi_1}{\partial \theta} \frac{\partial \psi_1}{\partial r} + \lambda_2 \frac{\partial \psi_2}{\partial \theta} \frac{\partial \psi_2}{\partial r} \right) d\theta. \quad (\text{C } 7)$$

This is the continuity equation for the pseudomomentum appearing in (5.2).

## Appendix D. Differential equation for topographic perturbations at $r > R_2$

Define the operators

$$D_1 = \partial_r^2 + \frac{1}{r} \partial_r - \frac{m^2}{r^2}, \quad D_2 = \partial_r^2 + \frac{1}{r} \partial_r - \frac{m^2}{r^2} - \frac{\Lambda^2}{\lambda_1 \lambda_2}, \quad (\text{D } 1)$$

which, according to the definitions of the BT and BC Green's functions (see Appendix B), satisfy

$$D_1 G_{\text{BT}}(r, r') = \delta(r - r'), \quad D_2 G_{\text{BC}}(r, r') = \delta(r - r'). \quad (\text{D } 2)$$

Define also

$$D_3 = \partial_r^2 - \frac{1}{r} \partial_r - \frac{m^2}{r^2}, \quad D_4 = \partial_r^2 - \frac{1}{r} \partial_r - \frac{m^2}{r^2} - \frac{\Lambda^2}{\lambda_1 \lambda_2}. \quad (\text{D } 3)$$

By applying the operator  $D_1 D_2$  to both sides of (6.9), using (3.18) and the identity

$$\int_{R_2}^{\infty} \delta^{(k)}(r) f(r) dr = (-1)^k \int_{R_2}^{\infty} \delta(r) f^{(k)}(r) dr, \quad (\text{D } 4)$$

we obtain the following fourth-order nonhomogeneous differential equation:

$$D_1 D_2 \xi(r) = - \frac{D_1 D_2 G_{22}(r, r_c)}{r_c} + \beta \lambda_2 D_4 \left[ \frac{1/r - 1/r_c}{\frac{\bar{V}_2(r)}{r} - \frac{\omega}{m}} \xi(r) \right] + \beta \lambda_1 D_3 \left[ \frac{1/r - 1/r_c}{\frac{\bar{V}_2(r)}{r} - \frac{\omega}{m}} \xi(r) \right], \quad (\text{D } 5)$$

where the source term is proportional to  $\delta(r - r_c)$  and its derivatives,

$$\begin{aligned} D_1 D_2 G_{22}(r, r_c) &= D_2 D_1 \lambda_2 G_{\text{BT}}(r, r_c) + D_1 D_1 \lambda_1 G_{\text{BC}}(r, r_c) \\ &= \lambda_2 D_2 \delta(r - r_c) + \lambda_1 D_1 \delta(r - r_c) \\ &= D_1 \delta(r - r_c) - \frac{\Lambda^2}{\lambda_1} \delta(r - r_c) \\ &= \delta''(r - r_c) + \frac{\delta'(r - r_c)}{r} - \left( \frac{m^2}{r^2} + \frac{\Lambda^2}{\lambda_1} \right) \delta(r - r_c). \end{aligned} \quad (\text{D } 6)$$

## Appendix E. Poles of response function

This appendix derives the types of poles of the response function  $\chi(r; r_0, \omega)$  defined by (6.30). Based on (6.30), for any  $\omega \neq m \bar{V}_2(r_0)/r_0$ ,  $\chi$  can be written as

$$\chi(r; r_0, \omega) = \frac{1}{\frac{\bar{V}_2(r_0)}{r_0} - \frac{\omega}{m}} \delta(r - r_0) + X(r; r_0, \omega), \quad (\text{E } 1)$$

where  $X(r, r_0; \omega)$  satisfies

$$\left(\frac{\bar{V}_2}{r} - \frac{\omega}{m}\right) X(r; r_0, \omega) - \beta \int_{R_2}^{\infty} \frac{G_{22}(r, r')}{r'} X(r'; r_0, \omega) dr' = \beta \frac{G_{22}(r, r_0)}{\frac{\bar{V}_2(r_0)}{r_0} - \frac{\omega r_0}{m}}. \quad (\text{E } 2)$$

By (E 1), a pole exists at  $\omega = m\bar{V}_2(r_0)/r_0$ . Equation (E 2) is singular at  $r = r_c = (m\bar{V}_2/r)^{-1}(\omega)$ , so we use the same ansatz as in (6.1),

$$X(r; r_0, \omega) = D(r_0, \omega) \delta\left(\frac{\bar{V}_2}{r} - \frac{\omega}{m}\right) - \frac{\beta}{\frac{\bar{V}_2}{r} - \frac{\omega}{m}} \xi(r; r_0, \omega), \quad (\text{E } 3)$$

where now the last term is not defined via the principal value and  $\xi$  is assumed to be a regular function of  $r$ . Substituting (E 3) into (E 2) gives

$$-\xi(r; r_0, \omega) - \frac{D(r_0, \omega)G_{22}(r, r_c)}{|(V(r)/r)'_{r_c}|r_c} + \int_{R_2}^{\infty} \frac{\beta G_{22}(r, r')}{\left(\frac{\bar{V}_2(r')}{r'} - \frac{\omega}{m}\right) r'} \xi(r'; r_0, \omega) dr' = \frac{G_{22}(r, r_0)}{\frac{\bar{V}_2(r_0)}{r_0} - \frac{\omega r_0}{m}}. \quad (\text{E } 4)$$

By substituting  $r = R$  into (E 4),  $D(r_0, \omega)$  can be expressed in terms of  $\xi$ :

$$D(r_0, \omega) = \frac{|(V(r)/r)'_{r_c}|r_c}{G_{22}(R, r_c)} \left[ \xi(R; r_0, \omega) - \int_{R_2}^{\infty} \frac{\beta G_{22}(R, r')}{\left(\frac{\bar{V}_2(r')}{r'} - \frac{\omega}{m}\right) r'} \xi(r'; r_0, \omega) dr' - \frac{G_{22}(R, r_0)}{\frac{\bar{V}_2(r_0)}{r_0} - \frac{\omega r_0}{m}} \right]. \quad (\text{E } 5)$$

Based on (E 5),  $D(r_0, \omega)$  has poles along the entire segment  $\mathcal{S}_1$  in addition to the poles of  $\xi(r; r_0, \omega)$ . The poles of  $\xi(r; r_0, \omega)$  also appear in the second term on the RHS of (E 3).

Upon solving (E 2) for  $\omega \notin \mathcal{S}_1$ , another class of singularities appear. Applying the operator  $D_1 D_2$  [where  $D_1$  and  $D_2$  are defined by (D 1)] to both sides of (E 2) and using (D 2) and (D 6) gives

$$\begin{aligned} D_1 D_2 \left[ \left(\frac{\bar{V}_2}{r} - \frac{\omega}{m}\right) X(r, r_0; \omega) \right] - \frac{\beta D_2 X(r, r_0; \omega)}{r} \\ = \beta \frac{\delta''(r - r_0) + \delta(r - r_0)/r_0 - (m^2/r^2 + \Lambda^2/\lambda_2)\delta(r - r_0)}{\frac{\bar{V}_2(r_0)}{r_0} - \frac{\omega r_0}{m}}. \end{aligned} \quad (\text{E } 6)$$

The solution to (E 6) may be found by first solving the homogeneous part, which gives four linearly independent solutions. Next, in any of the regions  $R_2 < r < r_0$  or  $r_0 < r$ , the solution is written as a linear combination of the four solutions with a total of eight constant coefficients (four for each region). In the asymptotic limit  $r \gg r_0$ , (E 2) is the same as (6.1), so only two solutions exist (i.e., the asymptotically BT and BC solutions, see §4.2). This leaves us with six constant coefficients. Applying (E 2) and its derivative at  $r = R_2$  produces two boundary conditions at  $r = R_2$  and four equations that match the solutions and their derivatives (up to the third-order derivative) on both sides of  $r = r_0$ . The matching conditions are determined from the delta terms on the RHS of (E 6). This makes six (nonhomogeneous) equations for the six unknown coefficients that we designate  $A_1, A_2, \dots, A_6$ . The equations can be recast to a standard matrix notation  $M(r_0, \omega)\mathbf{a} = \mathbf{b}$ , where  $\mathbf{a} = (A_1, \dots, A_6)$  and  $\mathbf{b} \neq \mathbf{0}$ . By Cramer's rule, the solutions are  $A_i = \frac{\det(M_i(r_0, \omega))}{\det(M(r_0, \omega))}$ , where  $M_i(r_0, \omega)$  is the matrix formed by replacing column  $i$  of  $M$  by the column vector  $\mathbf{b}$ . Therefore,  $X(r; r_0, \omega)$  is nonanalytic when the joint denominator of the coefficients,  $\det(M(r_0, \omega))$ , is zero. Since the determinant is a continuous function of  $\omega$ , its zeros constitute a discrete set of points.



## Appendix F. Sign of pseudomomentum via slope of dispersion curves

The proof here closely follows that of Ref. ?, which applies to single-layer shallow-water systems. Since the Rayleigh equation (3.4) is well defined for any  $m > 0$  (not necessarily an integer) here we treat  $m$  as a continuous variable, although in practice it must be an integer [see equation (3.3)]. Multiply the Rayleigh equation (3.4) by the complex conjugate of another solution  $\mathcal{Q}_1$  corresponding to a different mode number  $\tilde{m}$ ,

$$\left(\frac{\bar{V}_1(r)}{r} - \frac{\omega}{m}\right) \mathcal{Q}_1 \tilde{\mathcal{Q}}_1^* - \frac{\Phi_1}{r} \frac{d\bar{Q}_1}{dr} \tilde{\mathcal{Q}}_1^* = 0. \quad (\text{F } 1)$$

Multiply the Rayleigh equation (3.4) for the other solution by the complex conjugate of the first solution, and take the complex conjugate of the result:

$$\left(\frac{\bar{V}_1(r)}{r} - \frac{\tilde{\omega}^*}{\tilde{m}}\right) \tilde{\mathcal{Q}}_1^* \mathcal{Q}_1 - \frac{\tilde{\Phi}_1}{r} \frac{d\bar{Q}_1}{dr} \mathcal{Q}_1 = 0. \quad (\text{F } 2)$$

If the perturbation is stable, then  $\tilde{\omega}^* = \tilde{\omega}$ . The difference between the two equations is

$$0 = \left(\frac{\tilde{\omega}}{\tilde{m}} - \frac{\omega}{m}\right) \mathcal{Q}_1 \tilde{\mathcal{Q}}_1^* - \frac{1}{r} \frac{d\bar{Q}_1}{dr} (\Phi_1 \tilde{\mathcal{Q}}_1^* - \tilde{\Phi}_1^* \mathcal{Q}_1), \quad (\text{F } 3)$$

$$\begin{aligned} \left(\frac{\tilde{\omega}}{\tilde{m}} - \frac{\omega}{m}\right) \mathcal{Q}_1 \tilde{\mathcal{Q}}_1^* &= \frac{1}{r} \frac{d\bar{Q}_1}{dr} \left( \Phi_1 \frac{d^2 \tilde{\Phi}_1^*}{dr^2} + \frac{\Phi_1}{r} \frac{d\tilde{\Phi}_1^*}{dr} - \frac{\tilde{m}^2}{r^2} \Phi_1 \tilde{\Phi}_1^* - \frac{\Lambda^2}{\lambda_1} \Phi_1 (\tilde{\Phi}_1^* - \tilde{\Phi}_2^*) \right) \\ &\quad - \frac{1}{r} \frac{d\bar{Q}_1}{dr} \left( \Phi_1 \frac{d^2 \tilde{\Phi}_1^*}{dr^2} + \frac{\tilde{\Phi}_1^*}{r} \frac{d\Phi_1}{dr} - \frac{m^2}{r^2} \tilde{\Phi}_1^* \Phi_1 - \frac{\Lambda^2}{\lambda_1} \tilde{\Phi}_1^* (\Phi_1 - \Phi_2) \right). \end{aligned} \quad (\text{F } 4)$$

Similar equations can be written for the second layer,

$$\begin{aligned} \left(\frac{\tilde{\omega}}{\tilde{m}} - \frac{\omega}{m}\right) \mathcal{Q}_2 \tilde{\mathcal{Q}}_2^* &= \frac{1}{r} \frac{d\bar{Q}_2}{dr} \left( \Phi_2 \frac{d^2 \tilde{\Phi}_2^*}{dr^2} + \frac{\Phi_2}{r} \frac{d\tilde{\Phi}_2^*}{dr} - \frac{\tilde{m}^2}{r^2} \Phi_2 \tilde{\Phi}_2^* + \frac{\Lambda^2}{\lambda_2} \Phi_2 (\tilde{\Phi}_2^* - \tilde{\Phi}_1^*) \right) \\ &\quad - \frac{1}{r} \frac{d\bar{Q}_1}{dr} \left( \Phi_2 \frac{d^2 \tilde{\Phi}_2^*}{dr^2} + \frac{\tilde{\Phi}_2^*}{r} \frac{d\Phi_2}{dr} - \frac{m^2}{r^2} \tilde{\Phi}_2^* \Phi_2 + \frac{\Lambda^2}{\lambda_2} \tilde{\Phi}_2^* (\Phi_2 - \Phi_1) \right). \end{aligned} \quad (\text{F } 5)$$

Multiplying (F 4) by  $r\lambda_1$  and (F 5) by  $r\lambda_2$ , summing, and then integrating with respect to  $r$  gives, after taking the limit  $m \rightarrow \tilde{m}$ ,

$$\frac{d(\omega/m)}{dm} \int_R^\infty r \left( \lambda_1 \frac{d\bar{Q}_1}{dr} |d_1|^2 + \lambda_2 \frac{d\bar{Q}_2}{dr} |d_2|^2 \right) dr = -\frac{2m}{r^3} \int_R^\infty \lambda_1 |\Phi_1|^2 + \lambda_2 |\Phi_2|^2 dr. \quad (\text{F } 6)$$

Using the definition of the pseudomomentum (5.1) gives

$$\frac{d(\omega/m)}{dm} M = \frac{2m}{r^3} \int_R^\infty \lambda_1 |\Phi_1|^2 + \lambda_2 |\Phi_2|^2 dr. \quad (\text{F } 7)$$

The RHS is always positive, so  $M$  has the same sign as  $d(\omega/m)/dm$ .

## Appendix G. Rewriting eigenvalue equation in terms of contour-contour modes

The two CC perturbations types are denoted  $A$  and  $B$ ; each type corresponds to different perturbations of contours  $\alpha_1$  and  $\alpha_2$  and different frequencies. We write the perturbations in vector form for ease of notation, so the eigenvectors of the CC system

are

$$\begin{bmatrix} \alpha_1 \\ \alpha_2 \end{bmatrix}_A = \begin{bmatrix} \alpha_{1A} \\ \alpha_{2A} \end{bmatrix}, \quad \begin{bmatrix} \alpha_1 \\ \alpha_2 \end{bmatrix}_B = \begin{bmatrix} \alpha_{1B} \\ \alpha_{2B} \end{bmatrix}, \quad (\text{G } 1)$$

with eigenvalues  $\omega_a$  and  $\omega_b$ , respectively. A general perturbation of the contours of the full system can be written as

$$\begin{bmatrix} \alpha_1 \\ \alpha_2 \end{bmatrix} = a \begin{bmatrix} \alpha_{1A} \\ \alpha_{2A} \end{bmatrix} + b \begin{bmatrix} \alpha_{1B} \\ \alpha_{2B} \end{bmatrix}. \quad (\text{G } 2)$$

Plugging (G 2) into (4.9) and (4.10) and using the fact that the vectors in (G 1) are the CC eigenvectors, we get

$$\frac{\omega_a}{m} \alpha_{1A} a + \frac{\omega_b}{m} \alpha_{1B} b - \beta \Delta_1 \int_{R_2}^{\infty} \frac{G_{12}(R_1, r')}{r'} \eta(r') dr' = \frac{\omega}{m} (a \alpha_{1A} + b \alpha_{1B}), \quad (\text{G } 3)$$

$$\frac{\omega_a}{m} \alpha_{2A} a + \frac{\omega_b}{m} \alpha_{2B} b - \beta \Delta_2 \int_{R_2}^{\infty} \frac{G_{22}(R_2, r')}{r'} \eta(r') dr' = \frac{\omega}{m} (a \alpha_{2A} + b \alpha_{2B}). \quad (\text{G } 4)$$

Multiplying (G 3) by  $\alpha_{2B}$ , (G 4) by  $\alpha_{1B}$ , and subtracting gives

$$\begin{aligned} \frac{\omega}{m} (\alpha_{1A} \alpha_{2B} - \alpha_{2A} \alpha_{1B}) a &= \frac{\omega_a}{m} (\alpha_{1A} \alpha_{2B} - \alpha_{2A} \alpha_{1B}) a - \beta \Delta_1 \alpha_{2B} \int_{R_2}^{\infty} \frac{G_{12}(R_1, r')}{r'} \eta(r') dr' \\ &\quad + \beta \Delta_2 \alpha_{1B} \int_{R_2}^{\infty} \frac{G_{22}(R_2, r')}{r'} \eta(r') dr'. \end{aligned} \quad (\text{G } 5)$$

Multiplying (G 3) by  $\alpha_{2A}$  and (G 4) by  $\alpha_{1A}$  and subtracting gives

$$\begin{aligned} \frac{\omega}{m} (\alpha_{2A} \alpha_{1B} - \alpha_{1A} \alpha_{2B}) b &= \frac{\omega_b}{m} (\alpha_{2A} \alpha_{1B} - \alpha_{1A} \alpha_{2B}) b - \beta \Delta_1 \alpha_{2A} \int_{R_2}^{\infty} \frac{G_{12}(R_1, r')}{r'} \eta(r') dr' \\ &\quad + \beta \Delta_2 \alpha_{1A} \int_{R_2}^{\infty} \frac{G_{22}(R_2, r')}{r'} \eta(r') dr'. \end{aligned} \quad (\text{G } 6)$$

The third equation results from substituting (G 2) into (4.11),

$$\begin{aligned} \frac{\omega}{m} \eta(r) &= - \frac{G_{21}(r, R_1)}{R_1} (a \alpha_{1A} + b \alpha_{1B}) - \frac{G_{22}(r, R_2)}{R_2} (a \alpha_{2A} + b \alpha_{2B}) + \frac{\bar{V}_2(r)}{r} \eta(r) \\ &\quad - \beta \int_{R_2}^{\infty} \frac{G_{22}(r, r')}{r'} \eta(r') dr'. \end{aligned} \quad (\text{G } 7)$$